# Left-right supersymmetry and signals at the LHC

Katri Huitu Helsinki Institute of Physics



# **Outline:**

- Motivation
- Left-right symmetric extensions (LR) of the Standard Model
- LR models with supersymmetry (LRSUSY)
- Dark matter in LRSUSY models
- Signals of LRSUSY at the LHC



Electromagnetic and strong interactions conserve parity. Maybe parity is conserved at high energies also by weak interactions and broken dynamically Pati, Salam (1974); Mohapatra, Pati (1975); Senjanovic, Mohapatra (1975)

If weak interactions were parity conserving at a high scale, every left-handed particle would need a right-handed counterpart

right-handed neutrinos needed

After neutrinos were deemed massive, neutrino mass has also become a motivation.

Smallest gauge group where left-right symmetry can be implemented is  $SU(3)_{c} \times SU(2)_{I} \times SU(2)_{R} \times U(1)_{R-I}$  Davidson (1979), Mohapatra, Marshak (1980)

Gauge symmetry could originate from SO(10) or  $E_6$  unified model

Anomalies2020 / Katri Huitu

Add supersymmetry to the left-right model

 $R_{parity} = (-1)^{3(B-L)+2s}$  is an exact symmetry as B-L is a gauge symmetry;

if  $SU(2)_R \times U(1)_{B-L}$  is broken by L= $\pm 2$  triplet,  $R_{parity}$  remains unbroken also after symmetry breaking Mohapatra, PRD (1986); Font, Ibanez, Quevedo, PLB (1989); Martin PRD (1992)

LSP is a dark matter candidate

proton remains stable

There is no SUSY CP problem because of parity invariance Kuchimanchi, PRL (1996), Mohapatra, Rasin, PRL (1996)

Also strong CP problem may be solved

Beg, Tsao (1978); Mohapatra, Rasin, (1996); Babu, Dutta, Mohapatra (2002)

# Left-right symmetric extensions of the Standard Model; $SU(2)_L \times SU(2)_R \times U(1)_{B-L}$

**Fermions:** 

$$(Q_L)^i = \begin{pmatrix} u_L^i \\ d_L^i \end{pmatrix} = (\mathbf{3}, \mathbf{2}, \mathbf{1}, \frac{1}{3}), \ (Q_R)^i = \begin{pmatrix} u_R^i \\ d_R^i \end{pmatrix} = (\mathbf{3}, \mathbf{1}, \mathbf{2}, \frac{1}{3}),$$

$$(L_L)^i = \begin{pmatrix} \nu_L^i \\ \ell_L^i \end{pmatrix} = (\mathbf{1}, \mathbf{2}, \mathbf{1}, -1), (L_R)^i = \begin{pmatrix} \nu_R^i \\ \ell_R^i \end{pmatrix} = (\mathbf{1}, \mathbf{1}, \mathbf{2}, -1)$$

with L,R the left- and right-chiral components,  $P_{L,R} = (1 \pm \gamma_5)$ 

$$Q = I_{_{3L}} + I_{_{3R}} + \frac{B - L}{2}$$

Anomalies2020 / Katri Huitu

5

**Higgs sector:** need to break  $SU(2)_1 \times SU(2)_R \times U(1)_{B-1} \rightarrow SU(2)_1 \times U(1)_Y$ 

Break  $SU(2)_R \times U(1)_{B-1} \rightarrow U(1)_Y$  by a triplet:

$$\Delta_R = \begin{pmatrix} \Delta_R^+ / \sqrt{2} & \Delta_R^{++} \\ \Delta_R^0 & -\Delta_R^+ / \sqrt{2} \end{pmatrix} = (\mathbf{1}, \mathbf{1}, \mathbf{3}, 2); \ \langle \Delta_R^0 \rangle = v_R$$

For simplicity, assume that  $\Delta_L$  is decoupled Chang, Mohapatra, Parida, PRL (1984) Inert  $\Delta_L$  is also motivated by electroweak  $\rho$  parameter.

 $SU(2)_{I} \times U(1)_{V} \rightarrow U(1)_{em}$  by a bidoublet:

 $\Phi = \begin{pmatrix} \phi_1^0 & \phi_2^+ \\ \phi_1^- & \phi_2^0 \end{pmatrix} = (\mathbf{1}, \mathbf{2}, \mathbf{2}, \mathbf{0}); \ \langle \Phi \rangle = \begin{pmatrix} \kappa & 0 \\ 0 & \kappa' e^{i\alpha} \end{pmatrix}$  $K^0 - \overline{K^0}$  mixing constraints  $\propto \kappa \kappa' e^{i\alpha}$  $v_R \gg \kappa \gg \kappa'; \ \kappa^2 + \kappa'^2 = v_{SM}^2$ Beall, Bander, Soni (1982) Anomalies2020 / Katri Huitu

6

11.9.2020

Fermion masses can be found from:

$$\begin{aligned} \mathcal{L}_{Y} &= h^{a}_{q,ij} \overline{Q}_{L,i} \Phi_{a} Q_{R,j} + \widetilde{h}^{a}_{q,ij} \overline{Q}_{L,i} \widetilde{\Phi}_{a} Q_{R,j} + h^{a}_{\ell,ij} \overline{L}_{L,i} \Phi_{a} L_{R,j} + \widetilde{h}^{a}_{\ell,ij} \overline{L}_{L,i} \widetilde{\Phi}_{a} L_{R,j} \\ &+ f_{ij} L^{T}_{R,i} C i \tau_{2} \Delta_{R} L_{R,j} + h.c. \end{aligned}$$

with 
$$\widetilde{\Phi} = i\sigma_2 \Phi^* \sigma_2$$

Quark masses:

$$M_{u} = h_{u}\kappa + \tilde{h}_{u} e^{-i\alpha}\kappa'; M_{d} = h_{d}e^{i\alpha}\kappa' + \tilde{h}_{d}\kappa$$

Charged lepton masses  $M_{\ell} = h_{\ell} e^{i\alpha} \kappa' + \tilde{h}_{\ell} \kappa$ 



Neutrino masses  $M_{\nu} = \begin{pmatrix} 0 & m_D \\ m_D^T & M_N \end{pmatrix}; m_D = h_{\nu}\kappa + \tilde{h}_{\nu} e^{-i\alpha}\kappa'; M_N = fv_R$ 

Type-I seesaw, no neutrino Majorana mass:  $m_{\nu} = -m_D M_N^{-1} m_D^T$ 

Note: if  $\Delta_L$  included, one gets also Majorana mass for the light neutrinos, and thus type-II seesaw

Gell-Mann, Ramond, Slansky (1979); Yanagida (1979); Mohapatra, Senjanovic (1980); Schechter, Valle (1980)



11.9.2020

#### Gauge bosons, SU(2)<sub>L</sub> x SU(2)<sub>R</sub> x U(1)<sub>B-L</sub>

New gauge bosons connected to enlarged gauge group  $W_R$ , W,  $Z_R$ , Z,  $\gamma$ 

$$m_W^2 = \frac{g_L^2}{2} (\kappa^2 + \kappa'^2);$$
  $m_{W_R}^2 = g_R^2 v_R^2$ 

$$m_Z^2 = \frac{g_L^2}{2 \cos^2 \theta_W} (\kappa^2 + \kappa'^2); \qquad m_{Z_R}^2 = 2(g_R^2 + g_{BL}^2) v_R^2$$

Theoretically  $\frac{g_R}{g_L} \gtrsim 0.55$  Bhupal Dev, Mohapatra, Zhang, JHEP 05 (2016)

$$W - W_R$$
 mixing  $\approx \frac{g_L}{g_R} \frac{\kappa \kappa'}{v_R^2}$ 

~

PDG mass limits:  $m_{W_R} > 5.2$  TeV assuming  $m_{N_R} \ll m_{W_R}$ Note that  $m_{Z_R} > m_{W_R}$ 

Anomalies2020 / Katri Huitu



#### Tests of left-right model

Lepton number violating processes:

Ge <sup>Mo</sup> Neutrinoless double beta decay,  $0\nu\beta\beta$ Xe  $\langle m_{\beta\beta} \rangle$  (eV)  $10^{-1}$ W  $\mathbf{IH}$  $10^{-2}$  $\langle \overline{\nu}_{e} = \nu_{e}$ NH  $W^{-}$  $10^{-3}$ U  $10^{-2}$  $10^{-3}$  $10^{-1}$  $10^{-4}$ 50 100 150 m<sub>lightest</sub> (eV) Lepton flavour violation in А meson decays,  $\tau \rightarrow \mu, e\gamma, \mu \rightarrow e\gamma, \mu \leftrightarrow e$  –conversion,...

J. Engel et al, Rept.Prog.Phys (2017)

**∳**Ca

Se

Cdr Cdr



CMS, JHEP 05 (2018), arXiv: 1803.11116

$$W_R \to \mathrm{N}_R \ell \to \ell \ell j j$$

N<sub>R</sub> Majorana similar amounts of same sign and opposite sign dileptons Keung, Senjanovic, PRL 50 (1983) 1427





### Supersymmetric left-right symmetric model (LRSUSY)

#### The Higgses in LRSUSY

$$\begin{split} \Phi_{1} &= \begin{pmatrix} \Phi_{11}^{+} & \Phi_{11}^{0} \\ \Phi_{12}^{0} & \Phi_{12}^{-} \end{pmatrix} \sim (\mathbf{1}, \mathbf{2}, \mathbf{2}^{*}, 0), \qquad \Phi_{2} = \begin{pmatrix} \Phi_{21}^{+} & \Phi_{21}^{0} \\ \Phi_{22}^{0} & \Phi_{22}^{-} \end{pmatrix} \sim (\mathbf{1}, \mathbf{2}, \mathbf{2}^{*}, 0), \qquad \mathbf{2} = I_{3L} + I_{3R} + \frac{B - L}{2} \\ \Delta_{L} &= \begin{pmatrix} \frac{1}{\sqrt{2}} \Delta_{L}^{-} & \Delta_{L}^{0} \\ \Delta_{L}^{--} & -\frac{1}{\sqrt{2}} \Delta_{L}^{-} \end{pmatrix} \sim (\mathbf{1}, \mathbf{3}, \mathbf{1}, -2), \qquad \delta_{L} = \begin{pmatrix} \frac{1}{\sqrt{2}} \delta_{L}^{+} & \delta_{L}^{++} \\ \delta_{L}^{0} & -\frac{1}{\sqrt{2}} \delta_{L}^{+} \end{pmatrix} \sim (\mathbf{1}, \mathbf{3}, \mathbf{1}, 2) \\ \Delta_{R} &= \begin{pmatrix} \frac{1}{\sqrt{2}} \Delta_{R}^{-} & \Delta_{R}^{0} \\ \Delta_{R}^{--} & -\frac{1}{\sqrt{2}} \Delta_{R}^{-} \end{pmatrix} \sim (\mathbf{1}, \mathbf{1}, \mathbf{3}, -2), \qquad \delta_{R} = \begin{pmatrix} \frac{1}{\sqrt{2}} \delta_{R}^{+} & \delta_{R}^{++} \\ \delta_{R}^{0} & -\frac{1}{\sqrt{2}} \delta_{R}^{+} \end{pmatrix} \sim (\mathbf{1}, \mathbf{1}, \mathbf{3}, 2) \\ S \sim (\mathbf{1}, \mathbf{1}, \mathbf{1}, 0) , \end{split}$$

#### with VEVs

$$\langle \Phi_1 \rangle = \begin{pmatrix} 0 & 0\\ \frac{v_1}{\sqrt{2}} & 0 \end{pmatrix}, \quad \langle \Phi_2 \rangle = \begin{pmatrix} 0 & \frac{v_2}{\sqrt{2}}\\ 0 & 0 \end{pmatrix}, \quad \langle \Delta_R \rangle = \begin{pmatrix} 0 & \frac{v_R}{\sqrt{2}}\\ 0 & 0 \end{pmatrix}, \quad \langle \delta_R \rangle = \begin{pmatrix} 0 & 0\\ \frac{v_R'}{\sqrt{2}} & 0 \end{pmatrix}$$

 $\langle S \rangle = \frac{v_S}{\sqrt{2}} \,.$ 

Superpotential:

$$W = Q_L^T \boldsymbol{Y}_Q^{(i)} \Phi_i Q_R + L_L^T \boldsymbol{Y}_L^{(i)} \Phi_i L_R + L_L^T \boldsymbol{h}_{LL} \delta_L L_L + L_R^T \boldsymbol{h}_{RR} \Delta_R L_R + \lambda_L S \operatorname{Tr} [\Delta_L \delta_L] + \lambda_R S \operatorname{Tr} [\Delta_R \delta_R] + \lambda_3 S \operatorname{Tr} [\tau_2 \Phi_1^T \tau_2 \Phi_2] + \lambda_4 S \operatorname{Tr} [\tau_2 \Phi_1^T \tau_2 \Phi_1] + \lambda_5 S \operatorname{Tr} [\tau_2 \Phi_2^T \tau_2 \Phi_2] + \lambda_5 S^3 + \xi_F S$$

Problems with scalar potential:

to preserve electric charge and R-parity conservation in the vacuum, add

- non-renormalizable terms Mohapatra, Rasin, PRL 76 (1996), PRD 54 (1996)
- additional B-L triplets Aulakh, Melfo, Rasin, Senjanovic, PRD 58 (1998)
- include radiative corrections → scale < 15 TeV → hopes to find at the LHC Kuchimanchi, Mohapatra, PRD 48 (1993); Babu, Mohapatra, PLB 668 (2008); Basso, Fuks, Krauss, Porod, JHEP 07 (2015)</li>



#### **Constraints on spectrum**

Spectrum largely determined by  $\langle \Delta_R^0 \rangle \sim v_R \in [10, 15]$  TeV Basso, Fuks, Krauss, Porod, JHEP 07 (2015)

Scalar potential terms include Frank, Fuks, KH, Mondal, Rai, Waltari, PRD 101 (2020)

 $|\xi_{F} + \lambda_{L} \operatorname{Tr}[\Delta_{L} \delta_{L}] + \lambda_{R} \operatorname{Tr}[\Delta_{R} \delta_{R}] + \lambda_{3} \operatorname{Tr}[\tau_{2} \Phi_{1}^{T} \tau_{2} \Phi_{2}] + \lambda_{4} \operatorname{Tr}[\tau_{2} \Phi_{1}^{T} \tau_{2} \Phi_{1}]$  $+ \lambda_{5} \operatorname{Tr}[\tau_{2} \Phi_{2}^{T} \tau_{2} \Phi_{2}] + 3\lambda_{S} S^{2}|^{2}$ 

Thus, large  $v_R$  lifts the energy of the ground state. If sign( $\lambda_R$ )= - sign ( $\lambda_S$ ), can still find a parameter region, where m<sub>gauge bosons</sub> is not within reach of LHC or HE-LHC

 $W_R$  decays to superpartners around 20%  $\rightarrow m_{W_R} \ge 3.3$  TeV  $\rightarrow m_{Z_R} \ge 5.6$  TeV



...Constraints on spectrum

Impose  $m_h=125.1\pm0.3$  GeV tan  $\beta \ge 5$ , stop masses  $\approx$ a couple of TeV and stop mixings

Mass of the doubly charged Higgses generated at loop-level Babu, Mohapatra, PLB 668 (2008)  $m_{H^{\pm\pm}} > 350 \text{ GeV}$   $m_{H^{\pm\pm}}$  not much above 1 TeV fix  $BR(H^{\pm\pm} \rightarrow \tau^{\pm}\tau^{\pm})=92\%$ ,  $BR(H^{\pm\pm} \rightarrow \mu^{\pm}\mu^{\pm}/e^{\pm}e^{\pm})=4\%$ ,  $m_{H^{\pm\pm}} > 350 \text{ GeV}$  $(m_N)_{ij} = (h_{RR})_{ij} v_R \rightarrow m_{N_{\tau}} \approx 750 \text{ GeV}$ ,  $m_{N_e} \approx m_{N_{\mu}} \approx 150 \text{ GeV}$ 

 $m_{H^0_2} pprox m_{A^0_1} pprox m_{H^\pm_1}$  constrained by  $B_s o \mu \mu$  , which favors moderate tan eta

Other Higgs masses several TeV ( $\approx v_R$  or  $v_S$ ).

# Dark matter in LRSUSY

LRSUSY specific options: sneutrino and higgsino



Right-sneutrino mass is the only free parameter  $\rightarrow$  relic density determines  $m_{\tilde{\nu}_R} \approx 250 - 300 \text{ GeV}$ 

Frank, Fuks, KH, Rai, Waltari (2017); Chatterjee, Frank, Fuks, KH, Mondal, Rai, Waltari (2019)

# 2) SUSY partners of gauge bosons or Higgses can be the dark matter

Chatterjee, Frank, Fuks, KH, Mondal, Rai, Waltari (2019)

**Bidoublet higgsinos** form a nearly degenerate set of four neutralinos and two charginos

Coannihilations cannot be avoided, when the lightest higgsino is the LSP

The Planck-value for relic density is achieved with 750 GeV LSP higgsino (could be slightly decreased with further coannihilations, with e.g. sneutrino)

Note that in this case the spectrum is rather heavy and compressed difficult to detect

#### **Dominantly bino-like neutralino** with mass $m_h/2$ .

"Bino" refers to B-L –gaugino.

### Collider signals at the LHC

Particles light enough to be accessed experimentally? Consider  $W_R$  and  $H^{\pm\pm}$ Also  $A_1$ ,  $H_2$ ,  $H_1^{\pm}$  could be within LHC reach: almost degenerate with mass  $\gtrsim$  700 GeV

1) Robust signal of LR symmetry is  $W_R$  together with  $N_R$ Assume two different cases: dark matter from right sneutrino or higgsino

 $W_R$  production; decay to neutralinos, charginos  $BR(W_R \rightarrow jj) \approx 50 \%, BR(W_R \rightarrow N\ell) \approx 16 \%,$  $BR(W_R \rightarrow \tilde{\chi}\tilde{\chi}) \approx 22 \%,$  $\rightarrow$  Leptons and missing energy  $\rightarrow$  Electroweakino searches are relevant

Frank, Fuks, KH, Rai, Waltari in CERN Yellow Rep. Monogr. 7 (2019)



Signal Regions	Requirements	
SRA44	$   N_{\ell} = 3, N_{\text{OSSF}} \ge 1, N_{\tau} = 0, M_T > 160 \text{ GeV}, \not\!$	
SRD16	$N_{\ell} = 2, N_{\rm OS} = 1, N_{\rm SF} = 0, N_{\tau} = 1, M_{T2} > 100 \; {\rm GeV}, \not\!$	CMS, JHEP 03 (2018), 166



The most difficult scenario Frank, Fuks, KH, Mondal, Rai, Waltari, PRD 101 (2020) no.11, 115014

If W<sub>R</sub> too heavy to be detected, the  $H^{\pm\pm}$  may be the best (only) possibility for experimental detection (SR1: 4  $\tau_h$ , require 3 reconstructed  $\tau_h$ , SR2:  $2\tau_h 2\ell$ )



We studied detection of  $H^{\pm\pm}$  at HE-LHC (27 TeV, 15 ab<sup>-1</sup>) in a class of left-right supersymmetric scenarios favored by dark matter, with a relic density as measured by the Planck collaboration originating from the co-annihilations of multiple higgsino states of about 700 GeV

The only available decay mode is to leptons. As previously BR( $H^{\pm\pm} \rightarrow \tau^{\pm} \tau^{\pm}$ )=92%, BR( $H^{\pm\pm} \rightarrow \mu^{\pm} \mu^{\pm}$ )= BR( $H^{\pm\pm} \rightarrow e^{\pm} e^{\pm}$ )=4%

Here  $H^{\pm\pm}$  is Drell-Yan pair produced

#### We expect the detection of $H^{\pm\pm}$ at the HE-LHC up to around 1 TeV

	BP1	BP2	BP3
$m_{W_R}$	6550.5	7486.2	7486.2
$m_{Z'}$	10993.2	12563.6	12563.6
$m_{H_1^{\pm\pm}}$	875.0	1016.4	780.9
$m_{ ilde{\chi}_1^0}$	878.4	803.7	770.1
$m_{ ilde{\chi}^0_2}$	889.7	812.1	777.7
$m_{ ilde{\chi}^0_3}$	893.0	815.3	780.7
$m_{ ilde{\chi}_4^0}$	895.6	817.6	782.9
$m_{ ilde{\chi}_5^0}$	1032.2	1043.5	1048.0
$m_{ ilde{\chi}_1^\pm}$	886.4	809.3	775.1
$m_{ ilde{\chi}_2^\pm}$	893.5	815.7	781.2
$m_{\tilde{\chi}_1^{\pm\pm}}$	7413.2	8412.9	5619.2
$\mathrm{BR}(H_1^{\pm\pm} \to \tau^{\pm}\tau^{\pm})$	0.92	0.92	0.92
$\mathrm{BR}(H_1^{\pm\pm} \to \ell^\pm \ell^\pm)$	0.08	0.08	0.08





## Conclusions

- ✓ LRSUSY is a natural model for supersymmetric dark matter due to B-L –symmetry
- Excellent viable possibilities for supersymmetric dark matter, different from MSSM options
- ✓ Signals:
  - Detection of  $W_R$  together with  $N_R = LR$
  - Supersymmetric channel 
    LRSUSY
  - Discovery of  $H^{\pm\pm}$  seesaw, type I and type II
  - If  $W_R$  is too heavy to be detected, detection of  $H^{\pm\pm}$  is still possible
- ✓ With no discovery at 27 TeV, it is almost certain that something more than the minimal model would be needed

# Thank You.