

Ipsita Saha
Kavli IPMU

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東京大学 国際高等研究所 カブリ数物連携宇宙研究機構
KAVLI INSTITUTE FOR THE PHYSICS AND MATHEMATICS OF THE UNIVERSE

Improved $(g - 2)_\mu$ measurements and Supersymmetry

ANOMALIES 2020 – 11/09/2020

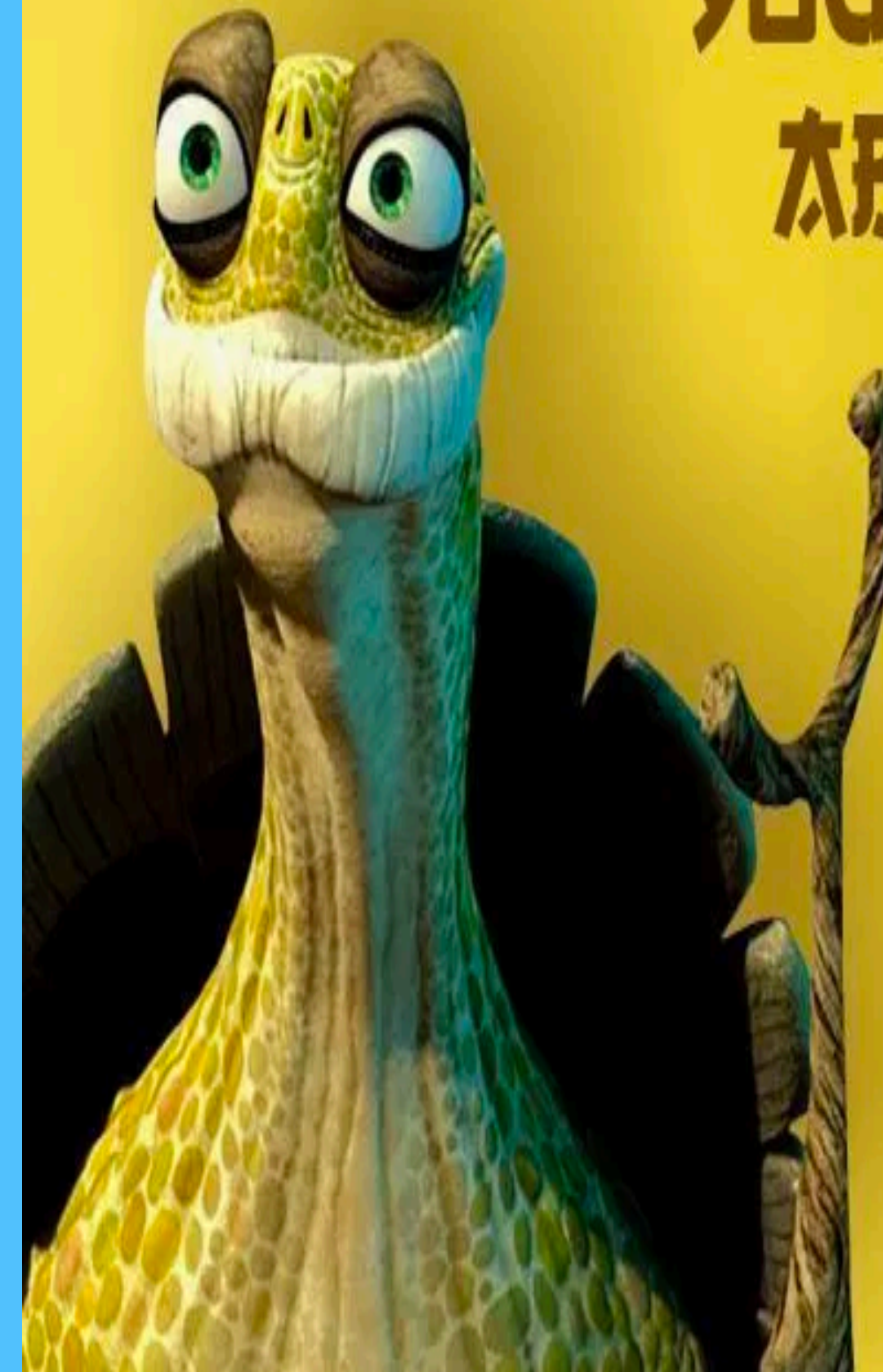
BASED ON : 2006.15157 WITH SVEN HEINEMEYER AND MANIMALA CHAKRABORTI

Possibilities :

- EW Sector may be the hiding key to new physics!
- Modest production cross-section, mass bounds from LHC rather weak.
- May show up elsewhere : Dark Matter experiments, $(g - 2)_\mu$ etc ..

MASTER OOGWAY QUOTES

SUSY, NO SUSY? **QUIT, DON'T QUIT?**
~~**NOODLES, DON'T NOODLES.**~~
**YOU ARE TOO CONCERNED
ABOUT WHAT WAS AND
WHAT WILL BE.**



MSSM Superpotential

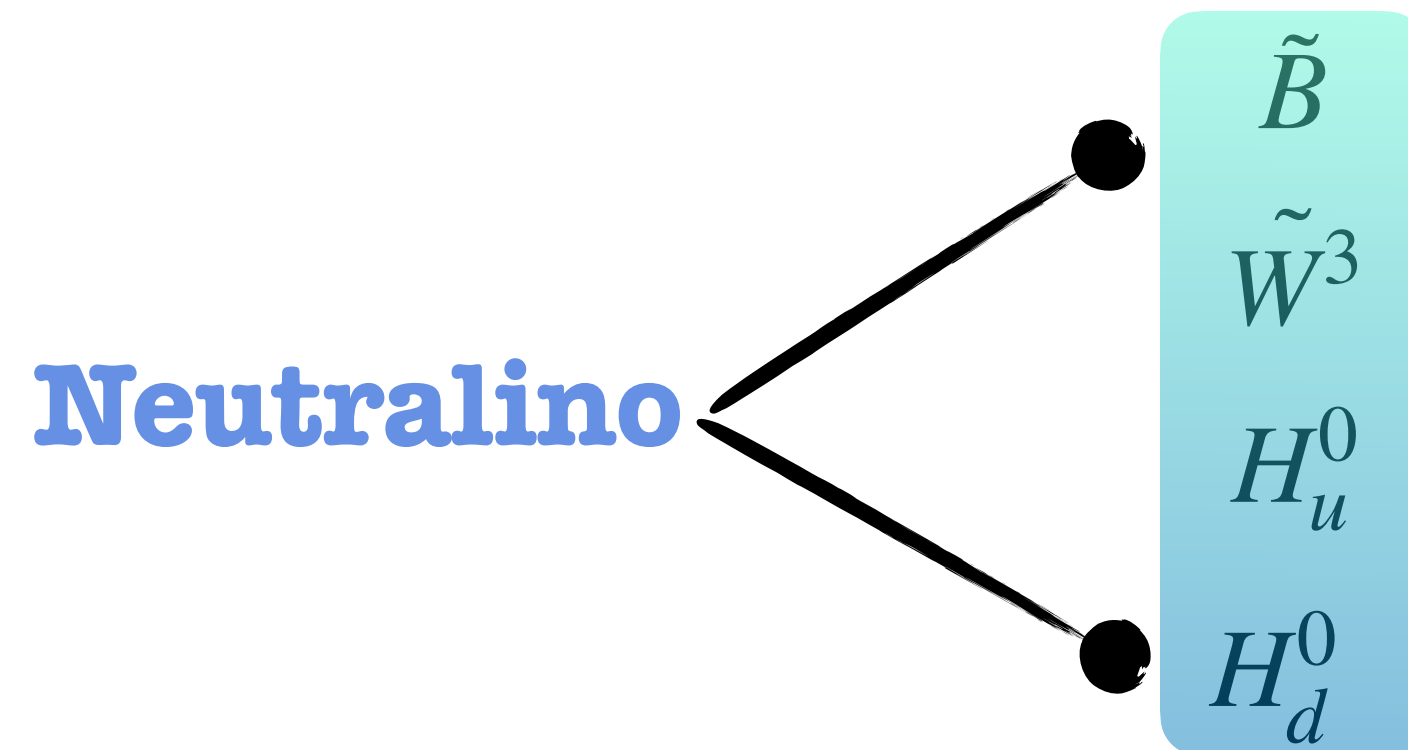
$$W_{\text{MSSM}} = \bar{u}Y_u QH_u - \bar{d}Y_d QH_d - \bar{e}Y_e LH_d + \mu H_u H_d$$

Soft Breaking Terms

$$\begin{aligned} \mathcal{L}_{\text{soft}}^{\text{MSSM}} = & -\frac{1}{2} (M_3 \tilde{g}\tilde{g} + M_2 \tilde{W}\tilde{W} + M_1 \tilde{B}\tilde{B} + c.c) \\ & - (\tilde{u} \mathbf{a}_u \tilde{Q}H_u - \tilde{d} \mathbf{a}_d \tilde{Q}H_d - \tilde{e} \mathbf{a}_e \tilde{L}H_d + c.c) \\ & - \tilde{Q}^\dagger \mathbf{m}_Q^2 \tilde{Q} - \tilde{L}^\dagger \mathbf{m}_L^2 \tilde{L} - \tilde{u} \mathbf{m}_u^2 \tilde{u}^\dagger - \tilde{d} \mathbf{m}_d^2 \tilde{d}^\dagger - \tilde{e} \mathbf{m}_e^2 \tilde{e}^\dagger \\ & - m_{H_u}^2 H_u^* H_u - m_{H_d}^2 H_d^* H_d - (bH_u H_d + c.c) \end{aligned}$$

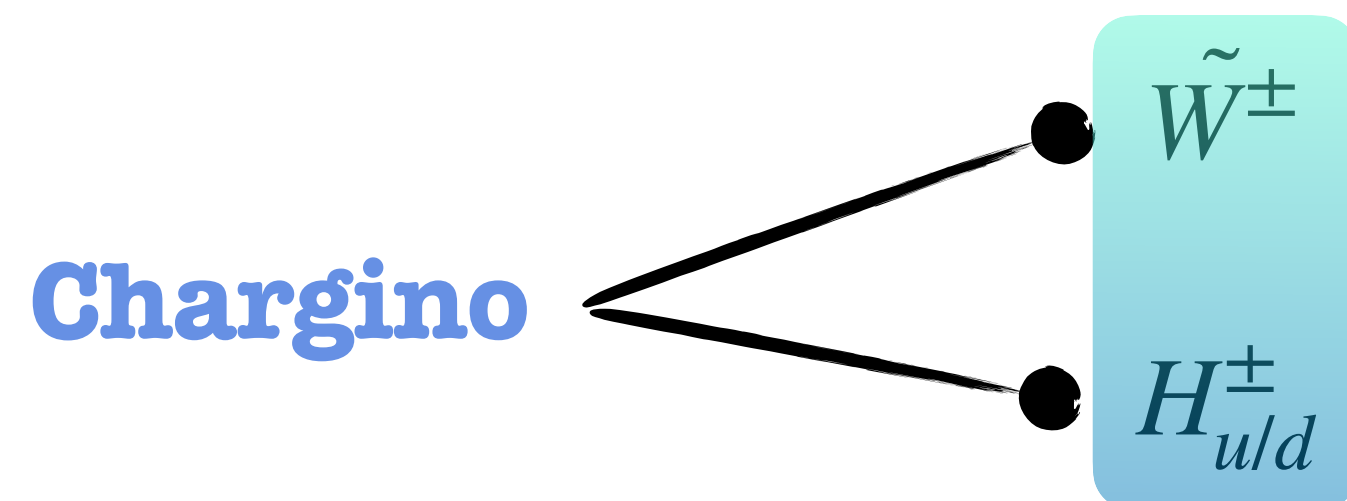
EW Gauginos

Masses and mixing are determined by U(1) and SU(2) gaugino masses M_1 , M_2 and Higgs mass parameter μ .



Neutralino Mass Matrix

$$M_N = \begin{pmatrix} M_1 & 0 & -M_Z c_\beta s_W & M_Z s_\beta s_W \\ 0 & M_2 & M_Z c_\beta c_W & -M_Z s_\beta c_W \\ -M_Z c_\beta s_W & M_Z c_\beta c_W & 0 & -\mu \\ M_Z s_\beta s_W & -M_Z s_\beta c_W & -\mu & 0 \end{pmatrix}$$



Chargino Mass Matrix

$$M_C = \begin{pmatrix} M_2 & \sqrt{2}M_W c_\beta \\ \sqrt{2}M_W s_\beta & \mu \end{pmatrix}$$

FOUR PARAMETERS



$M_1, M_2, \mu, \tan \beta$



Sleptons

Slepton Mass Matrix

$$M_{\tilde{L}}^2 = \begin{pmatrix} m_l^2 + m_{LL}^2 & m_l X_l \\ m_l X_l & m_l^2 + m_{RR}^2 \end{pmatrix}$$

PARAMETERS



$M_1, M_2, \mu, \tan \beta, m_{\tilde{L}}, m_{\tilde{R}}$

$$m_{LL}^2 = m_{\tilde{L}}^2 + (I_l^{3L} - Q_f s_w^2) M_z^2 c_\beta^2$$

$$m_{RR}^2 = m_{\tilde{R}}^2 + Q_f s_w^2 M_z^2 c_\beta^2$$

$$X_l = A_l - \mu (\tan \beta)^{2I_l^{3L}}$$

Constraints

Proper recasting is important.

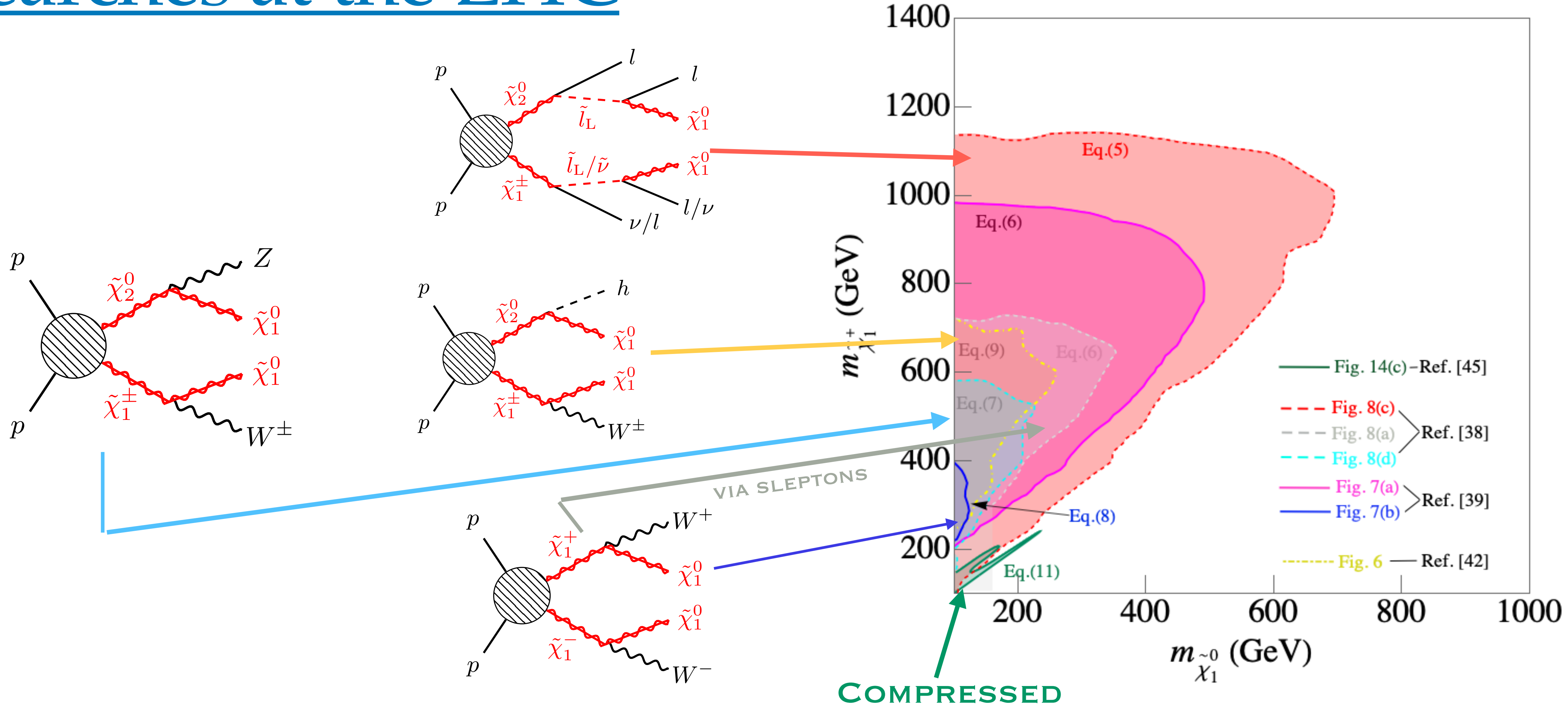
Direct Searches at LHC

- LHC searches restricted to **simplified models**
 - sparticles except those relevant to the signal are taken to be decoupled.
- $\tilde{\chi}_1^\pm$ and $\tilde{\chi}_2^0$ are taken to be mass-degenerate and purely wino. $\tilde{\chi}_1^0$ is assumed to be purely bino.
- All three generations of sleptons and sneutrinos are assumed mass degenerate. In MSSM:
$$m_{\tilde{\nu}}^2 = m_{\tilde{l}}^2 + \frac{1}{2}m_Z^2 \cos 2\beta$$
- Heavier gauginos $\tilde{\chi}_3^0, \tilde{\chi}_4^0, \tilde{\chi}_2^\pm$ assumed to be decoupled.
- No sensitivity to parameters like $\tan \beta$.

Indirect Constraints

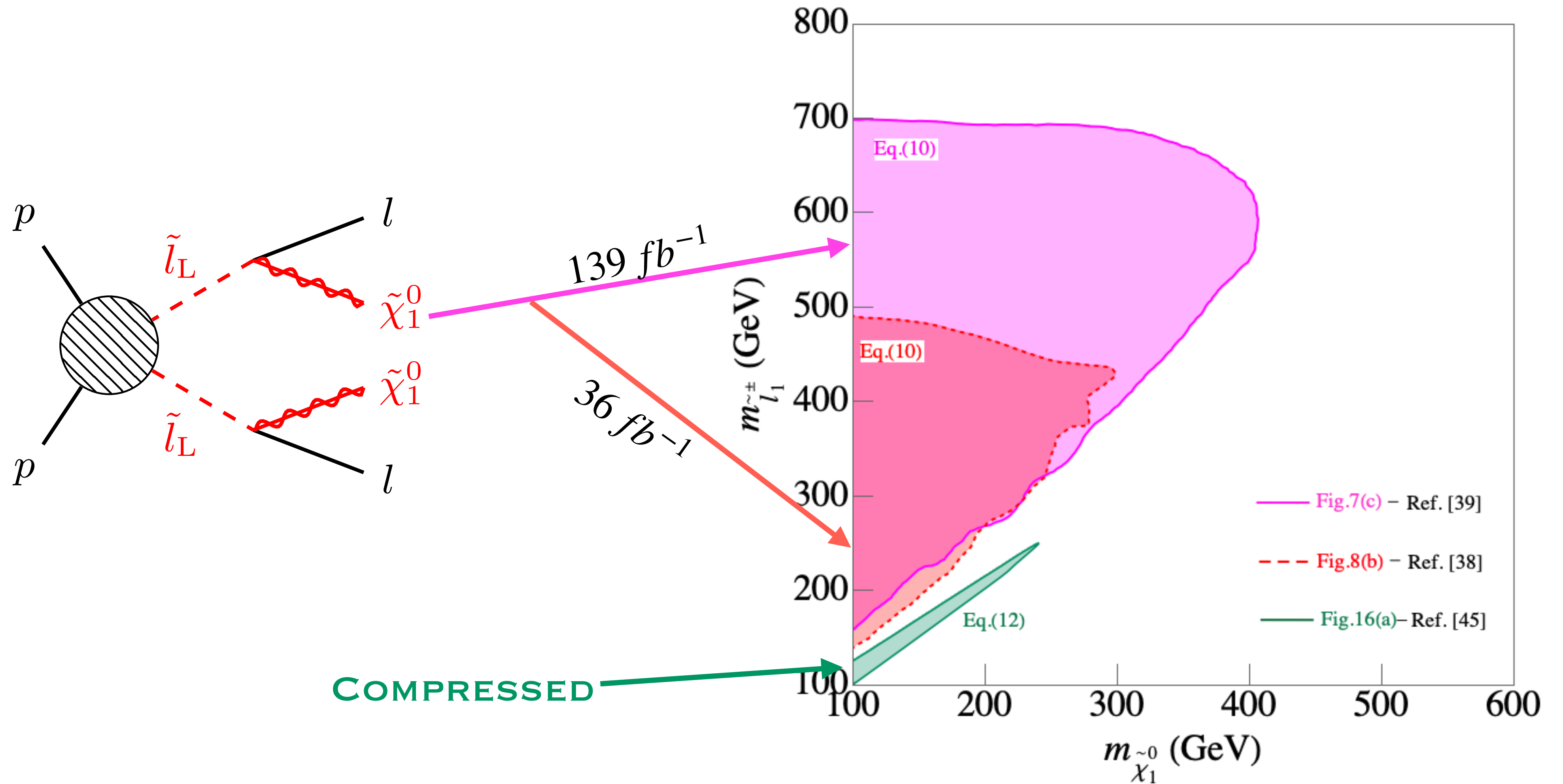
- Muon (g-2).
- WMAP/PLANCK relic density.
- Spin independent direct detection data from XENON/LUX.
- Indirect detection constraints of dark matter.

Searches at the LHC



Proper recasting is important. \longrightarrow CheckMATE

Searches at the LHC

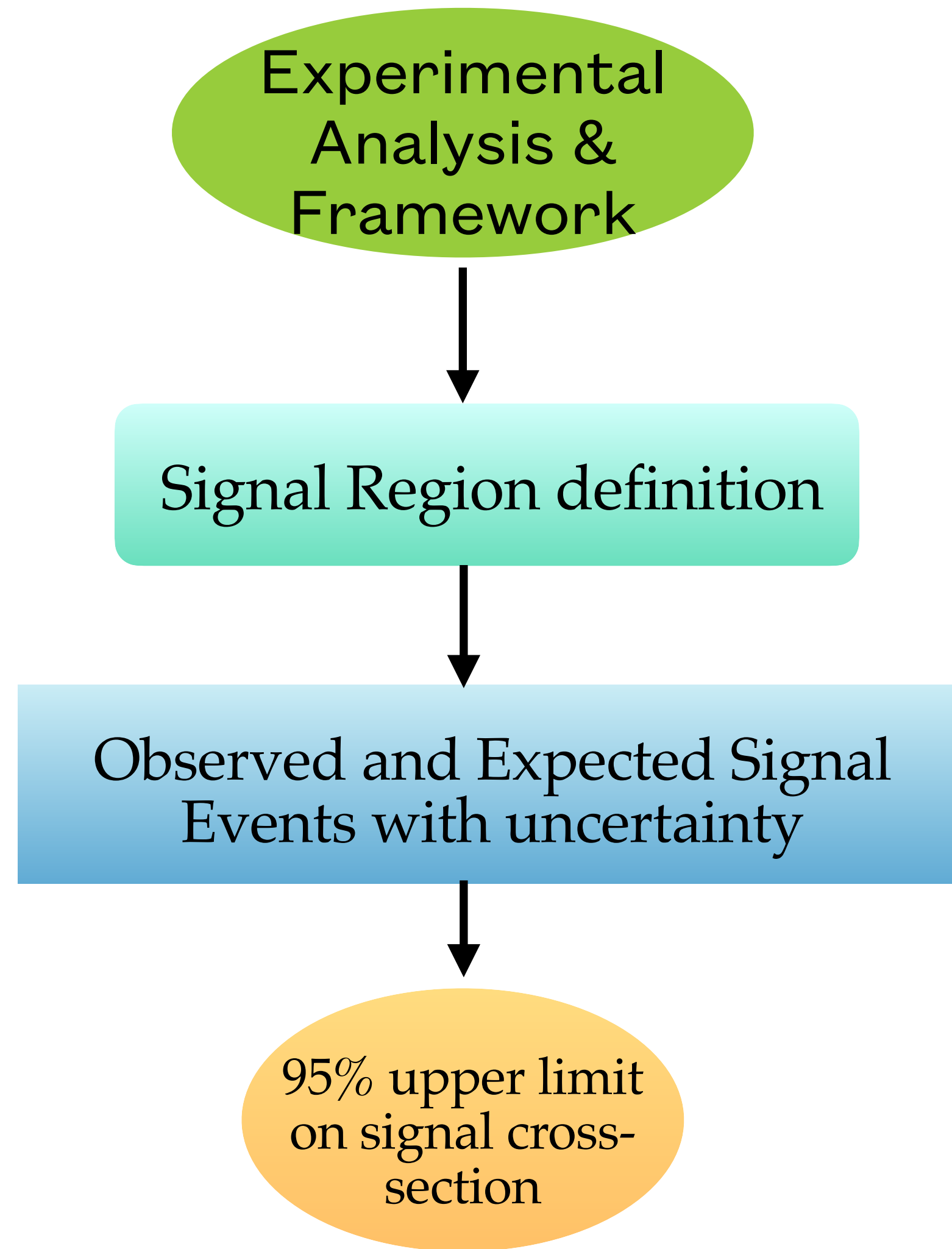


Proper recasting is important. \longrightarrow CheckMATE

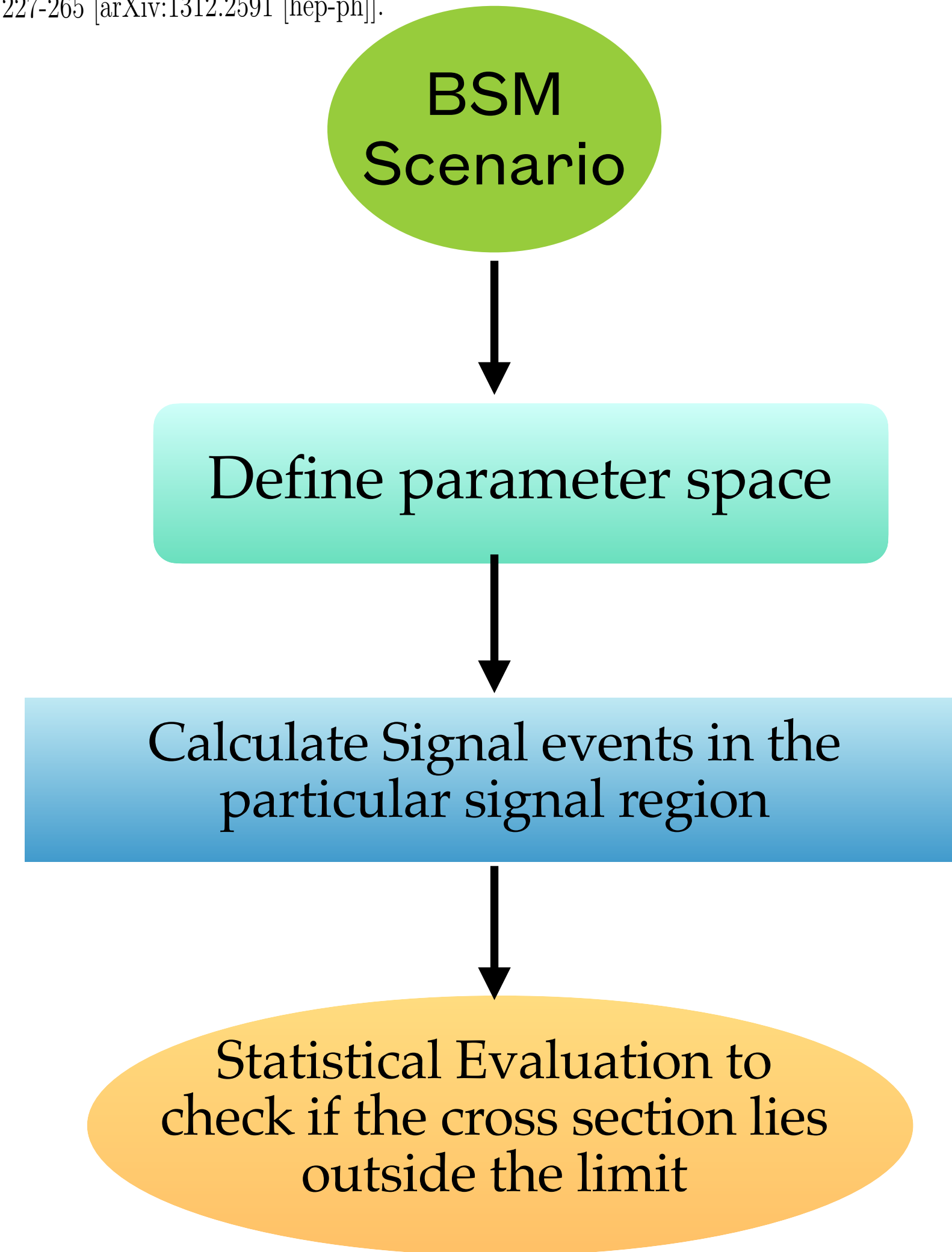
CheckMATE in brief

J. S. Kim, D. Schmeier, J. Tattersall and K. Rolbiecki, *Comput. Phys. Commun.* **196** (2015), 535-562 [arXiv:1503.01123 [hep-ph]].

M. Drees, H. Dreiner, D. Schmeier, J. Tattersall and J. S. Kim, *Comput. Phys. Commun.* **187** (2015), 227-265 [arXiv:1312.2591 [hep-ph]].



Input for Implementation of new analysis



Model parameter test

Muon (g-2)

- Currently large discrepancy from the SM $> (3\sigma)$.

$$a_{\mu}^{exp} - a_{\mu}^{SM} = (28.02 \pm 7.37) \times 10^{-10}$$

Kashavarzi, Nomura, Teubner '19

- Assuming upcoming Fermilab Run-I result to have the same central value and reduced exp uncertainty combined data corresponds to 5.4σ discrepancy.

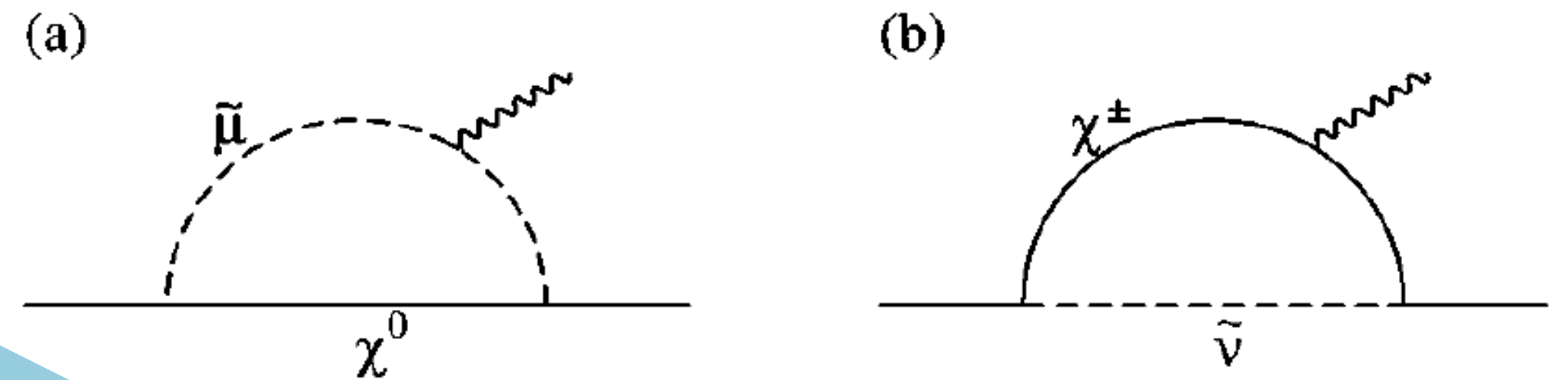
$$a_{\mu}^{exp} - a_{\mu}^{SM} = (28.02 \pm 5.2) \times 10^{-10}$$

- New “World average” appeared



modest impact on our analysis

Aoyama *et al* '20

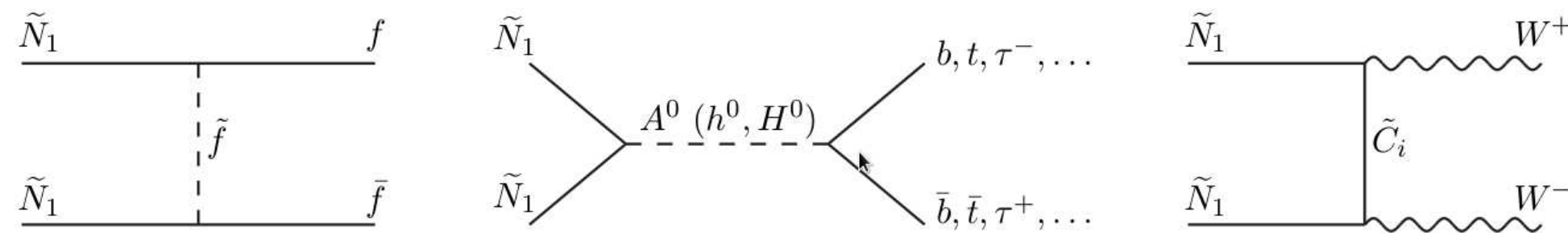


- SUSY contributions from Chargino-Sneutrino and Smuon-Neutralino loop
- Contribution $\sim \tan\beta$, can reconcile the anomaly

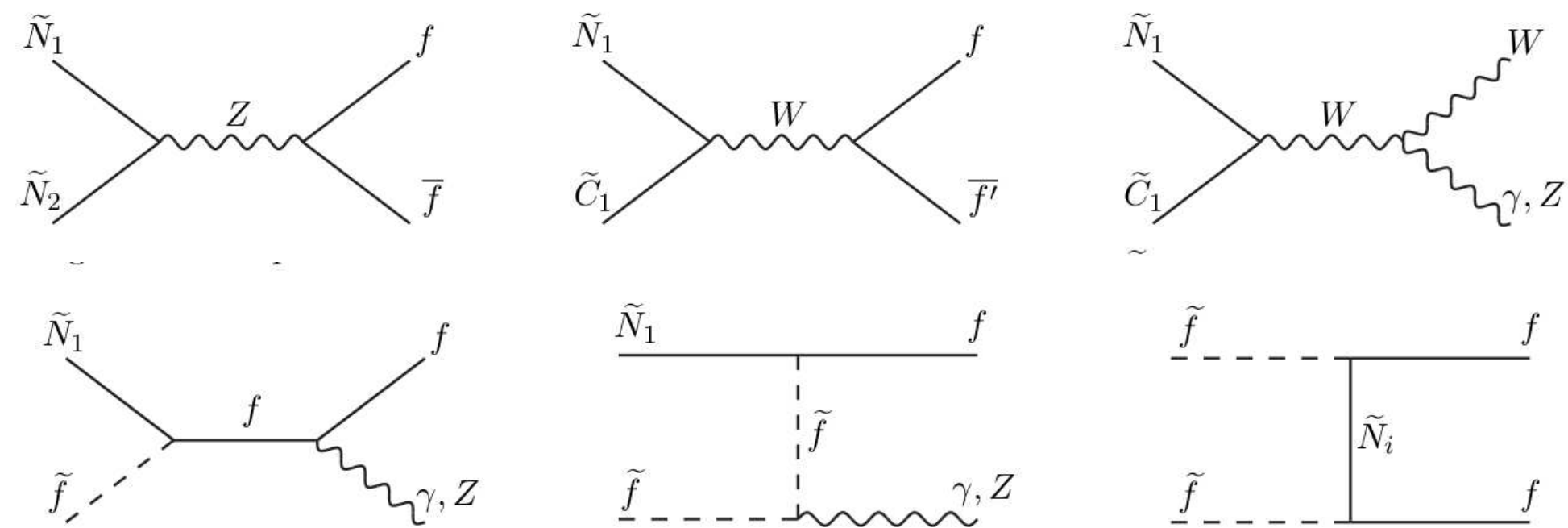
DM Constraints

Relic Density

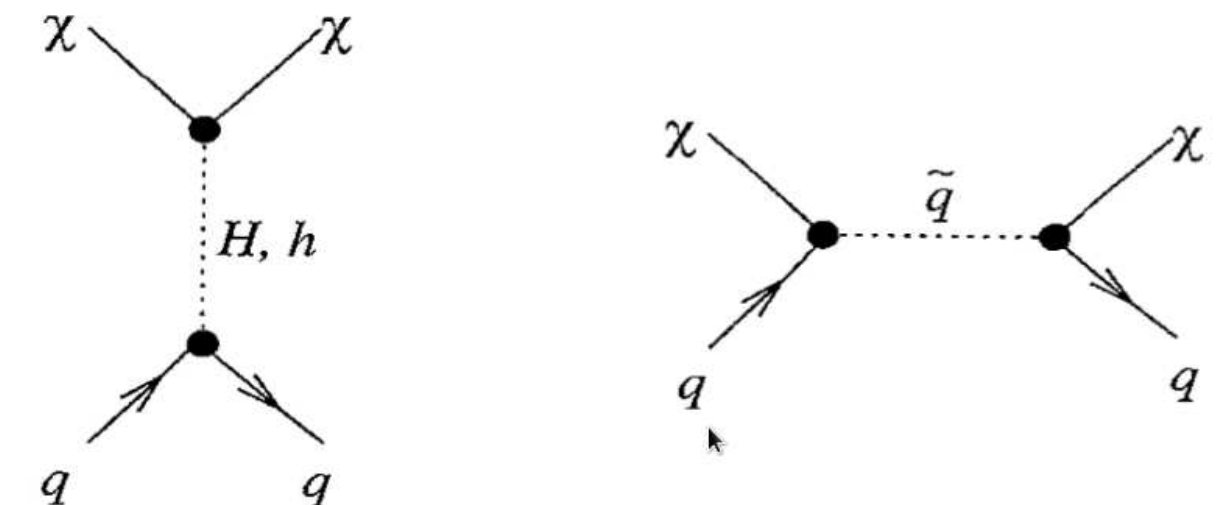
Some annihilation channels that could give right relic density :



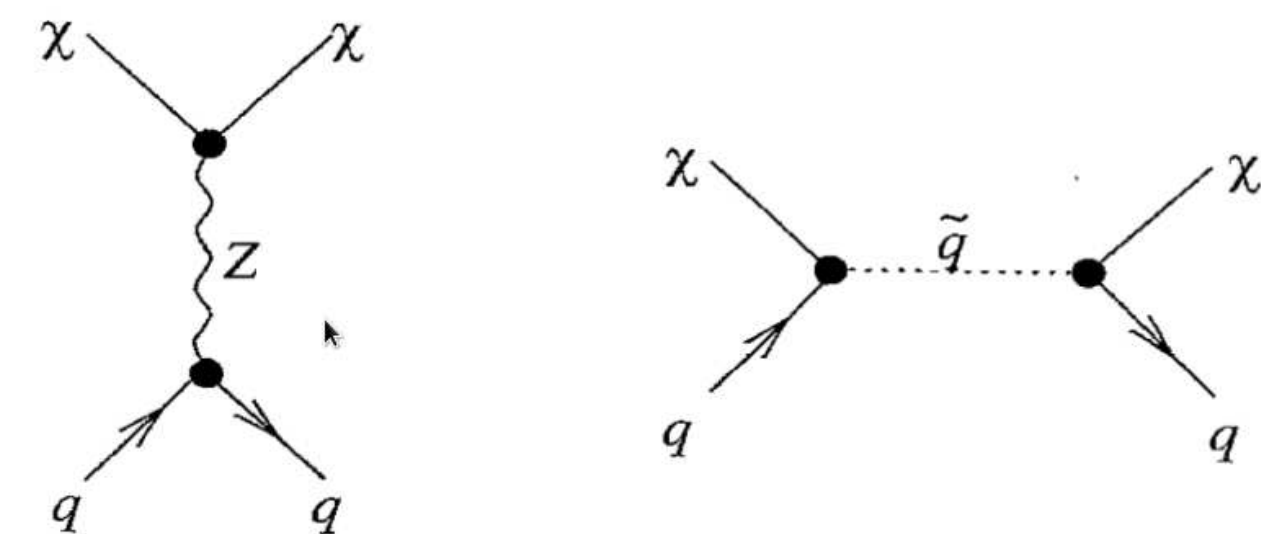
There can be coannihilations with sparticles of slightly heavier masses:



Direct Detection



Diagrams contributing to SI interactions



Diagrams contributing to SD interactions

A well-tempered bino-wino or bino-higgsino LSP is favorable for chargino co-annihilation while a bino dominated LSP will work for slepton co-annihilation.

Parameter Scanning

Chargino co-annihilation region:

$$100 \text{ GeV} \leq M_1 \leq 1 \text{ TeV}, \quad M_1 \leq M_2 \leq 1.1M_1, \\ 1.1M_1 \leq \mu \leq 10M_1, \quad 5 \leq \tan \beta \leq 60, \\ 100 \text{ GeV} \leq m_{\tilde{l}_L} \leq 1 \text{ TeV}, \quad m_{\tilde{l}_R} = m_{\tilde{l}_L}.$$

Slepton co-annihilation region:

Case-L: SU(2) doublet

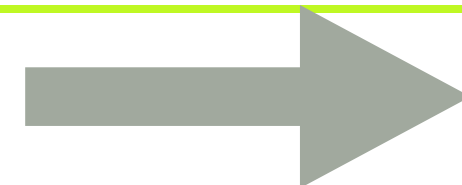
$$100 \text{ GeV} \leq M_1 \leq 1 \text{ TeV}, \quad M_1 \leq M_2 \leq 10M_1, \\ 1.1M_1 \leq \mu \leq 10M_1, \quad 5 \leq \tan \beta \leq 60, \\ M_1 \text{ GeV} \leq m_{\tilde{l}_L} \leq 1.2M_1, \quad M_1 \leq m_{\tilde{l}_R} \leq 10M_1.$$

Case-R: SU(2) singlet

$$100 \text{ GeV} \leq M_1 \leq 1 \text{ TeV}, \quad M_1 \leq M_2 \leq 10M_1, \\ 1.1M_1 \leq \mu \leq 10M_1, \quad 5 \leq \tan \beta \leq 60, \\ M_1 \text{ GeV} \leq m_{\tilde{l}_R} \leq 1.2M_1, \quad M_1 \leq m_{\tilde{l}_L} \leq 10M_1.$$

MC, S.Heinemeyer, I.Saha 2006.15157

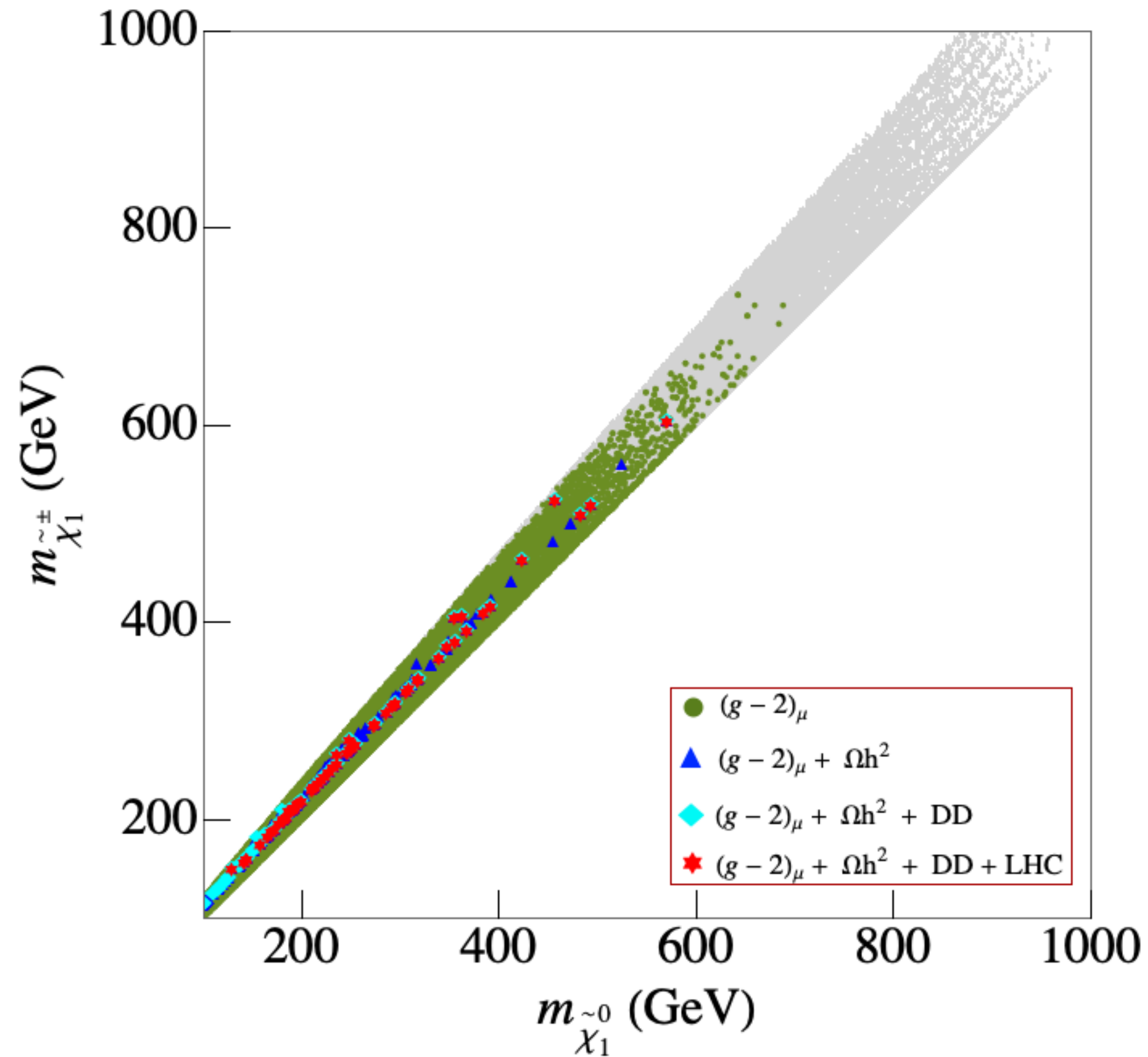
Packages used



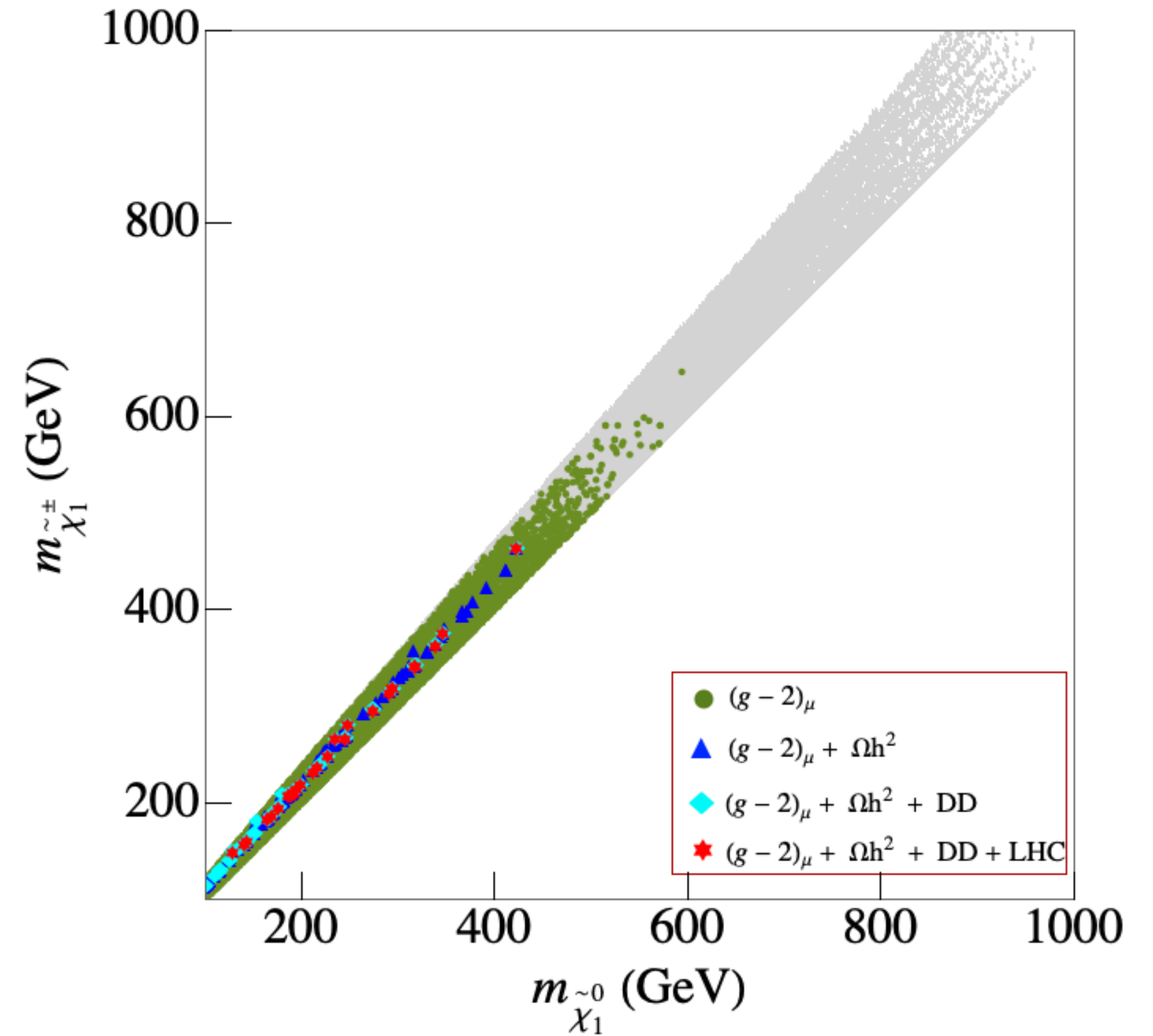
SuSpect, SUSYHIT, GM2Calc, micrOMEGAs, CheckMATE etc.

Chargino Co-annihilation

Current $(g - 2)_\mu$ limit

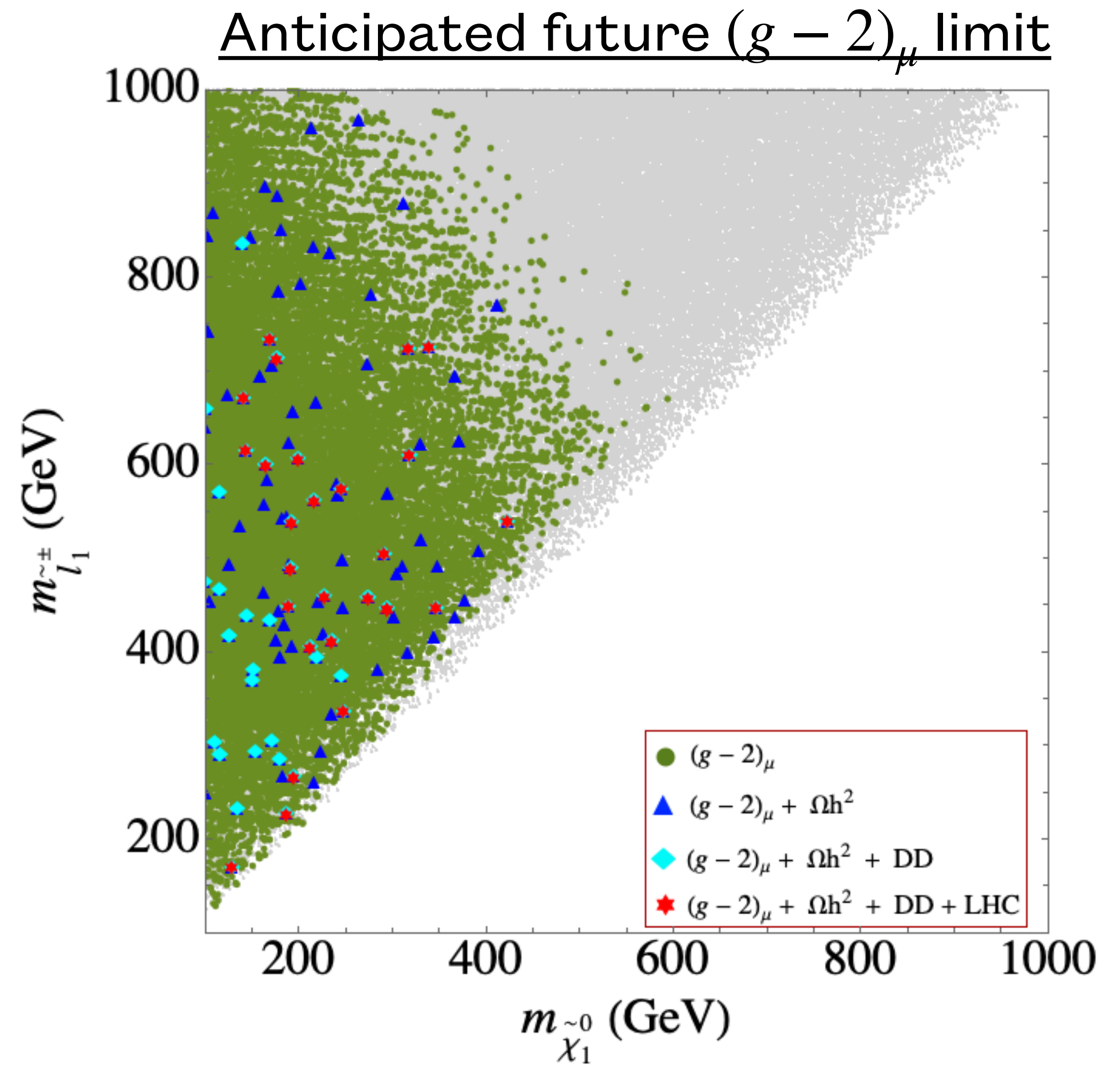
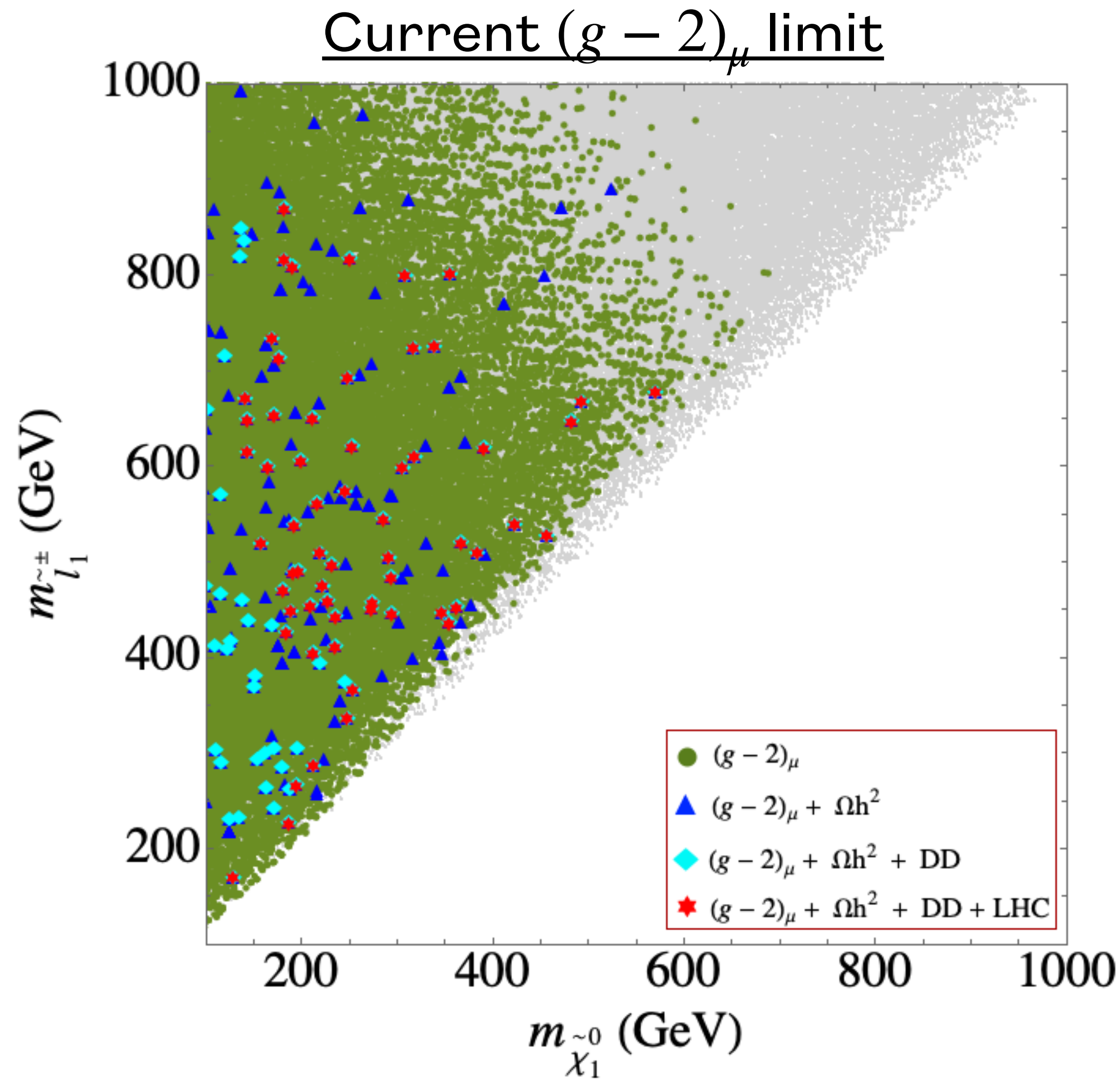


Anticipated future $(g - 2)_\mu$ limit



Upper and lower bounds from $(g - 2)_\mu$ and LHC searches (for compressed spectrum)

Chargino Co-annihilation

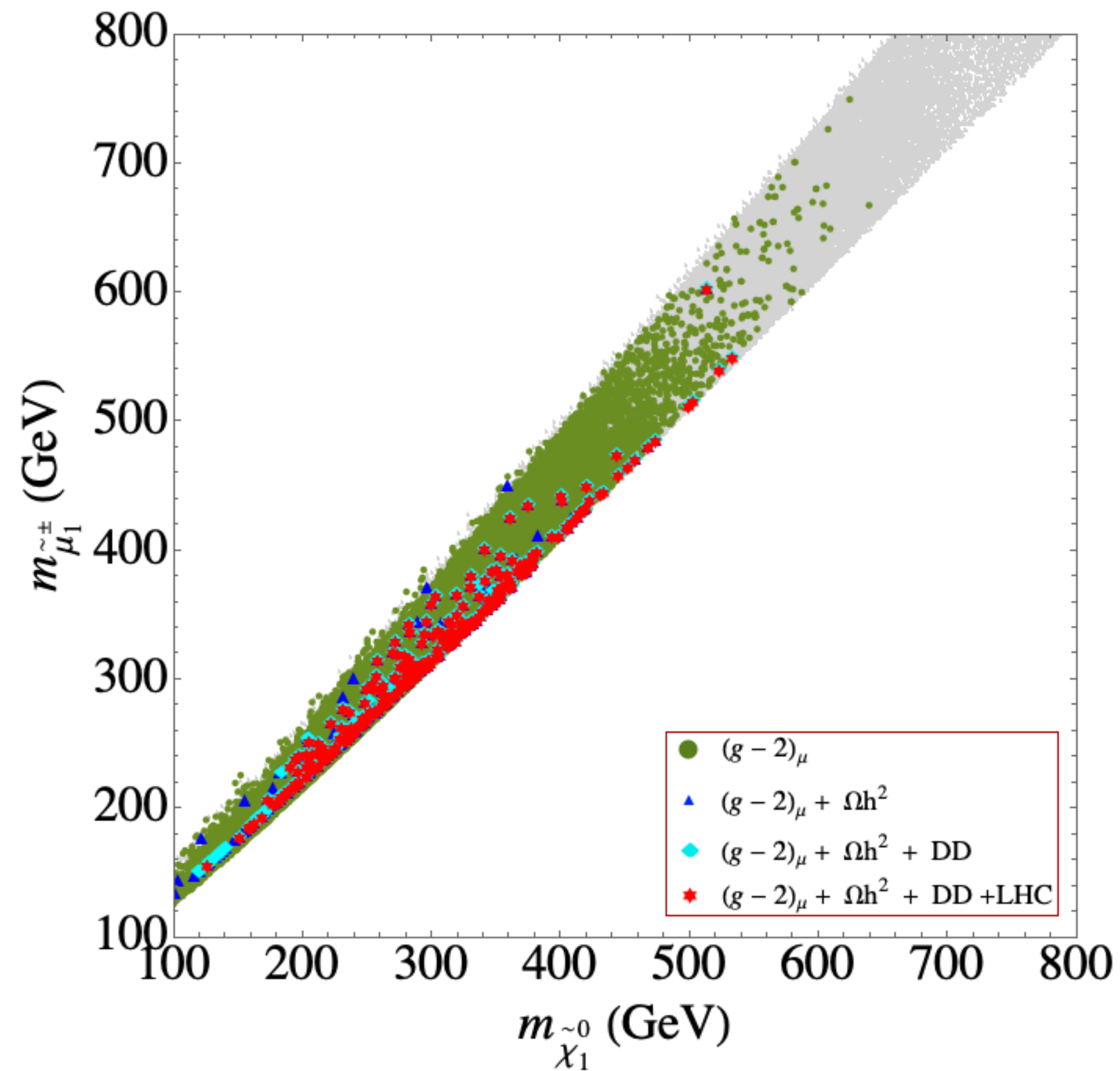


Slepton-pair production $\rightarrow (2l + \text{missing } E_T)$ provides important search channel.

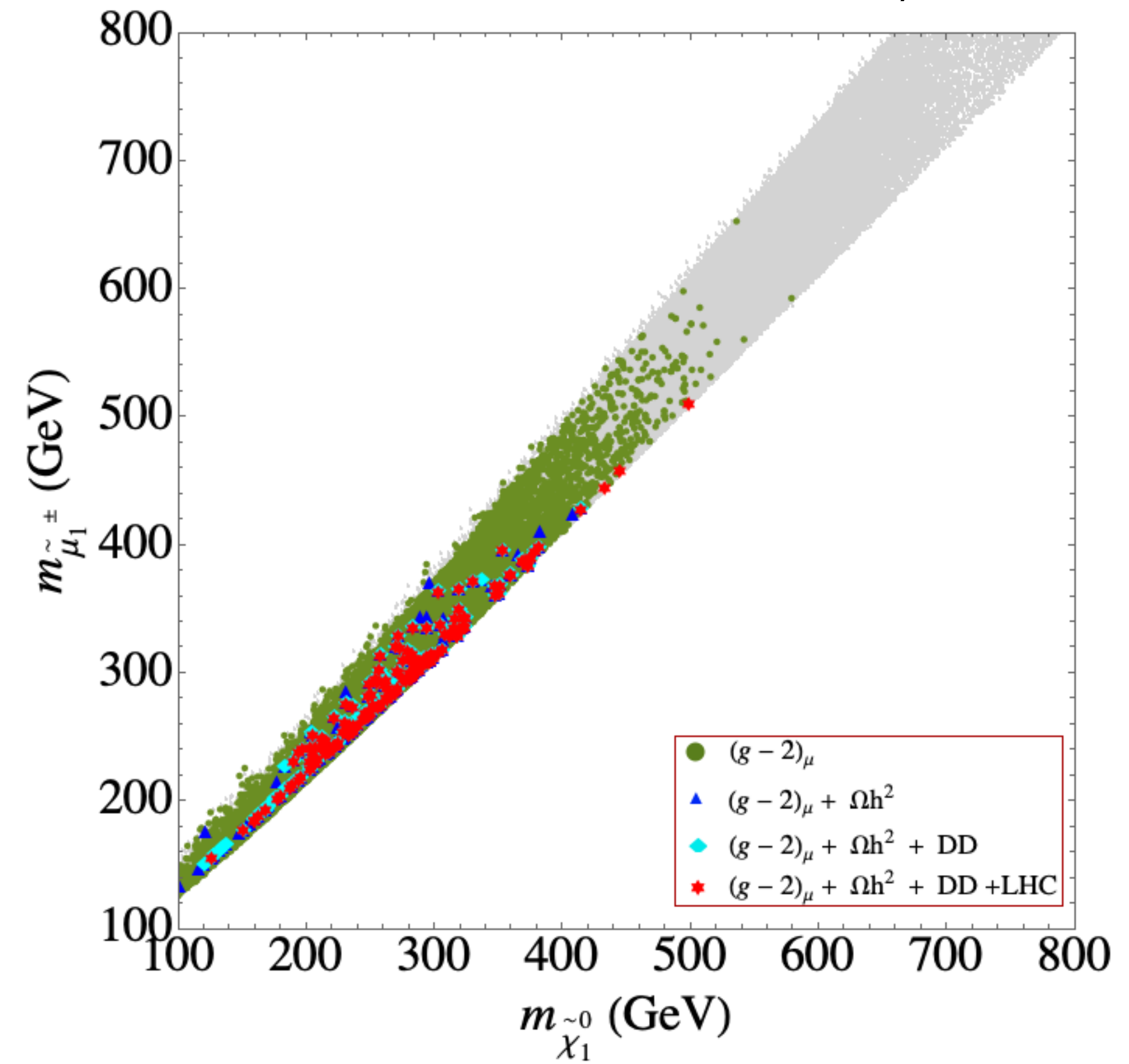
Right-sleptons are significantly heavy, Considerable BR for $\tilde{e}_L(\tilde{\mu}_L) \rightarrow \tilde{\chi}_1^\pm e(\mu)$ \longrightarrow Less no. of signal leptons.

Slepton Co-annihilation: Case-L

Current $(g - 2)_\mu$ limit



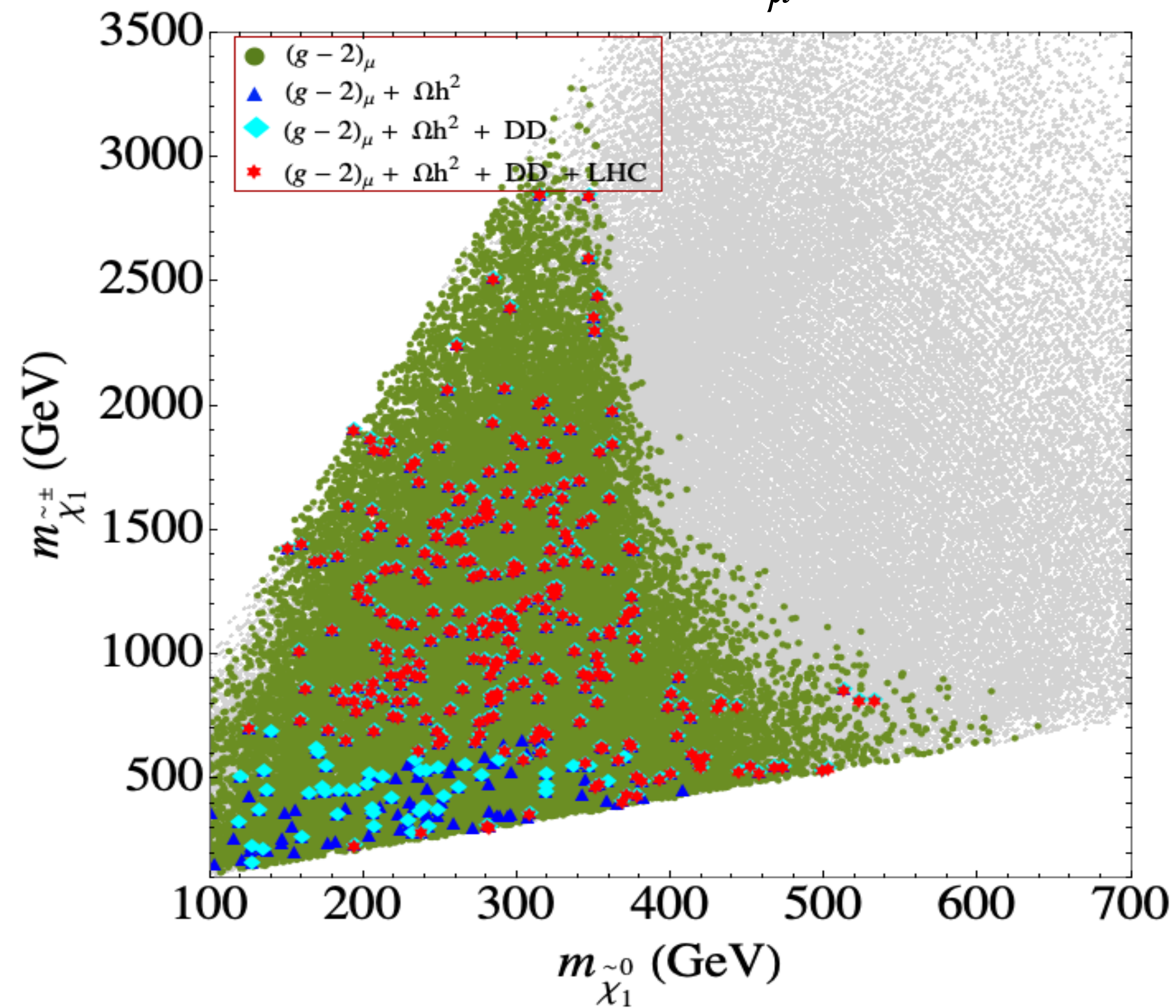
Anticipated future $(g - 2)_\mu$ limit



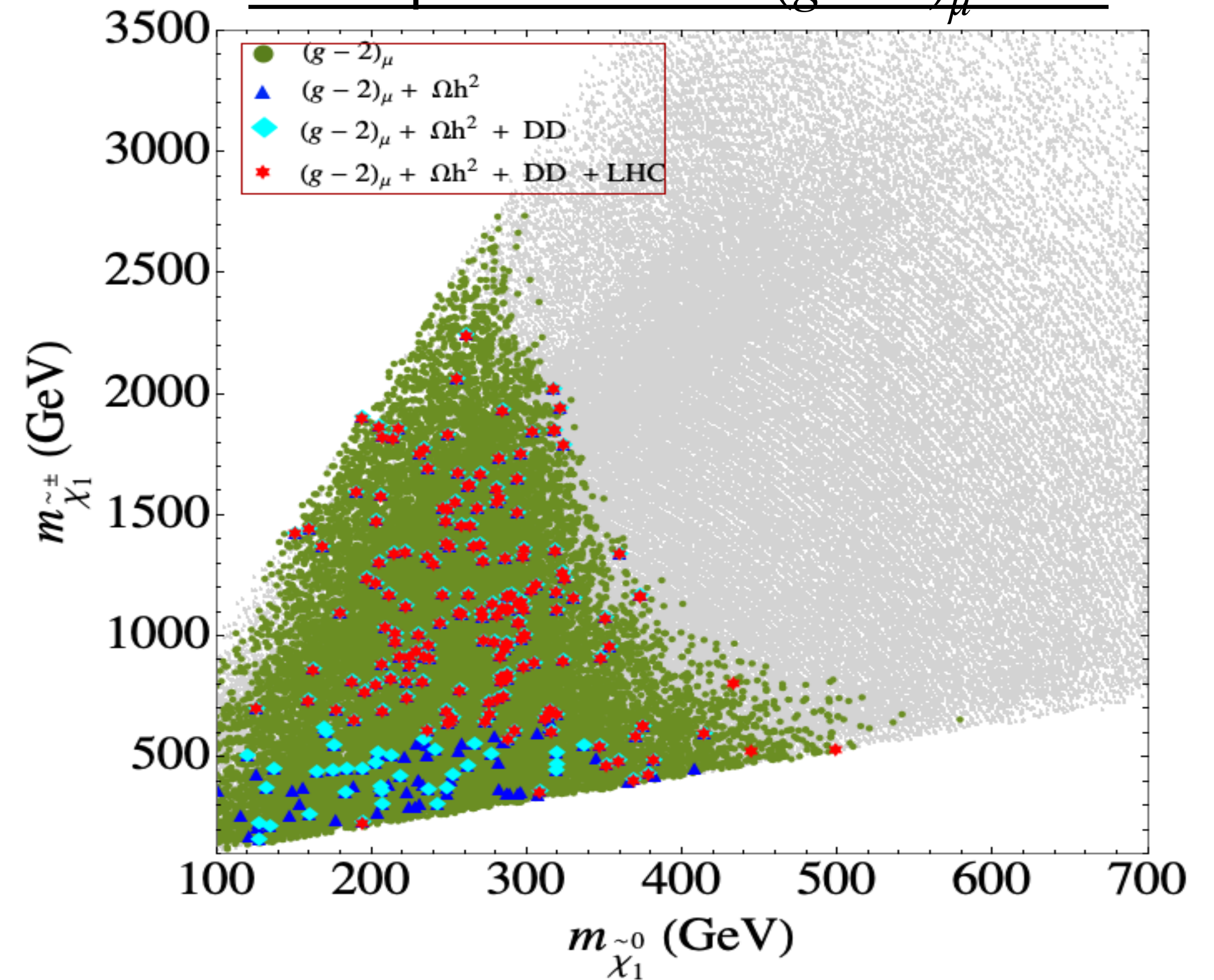
The left-sleptons and sneutrinos are close in mass to the LSP \longrightarrow stau not far away

Slepton Co-annihilation: Case-L

Current $(g - 2)_\mu$ limit



Anticipated future $(g - 2)_\mu$ limit

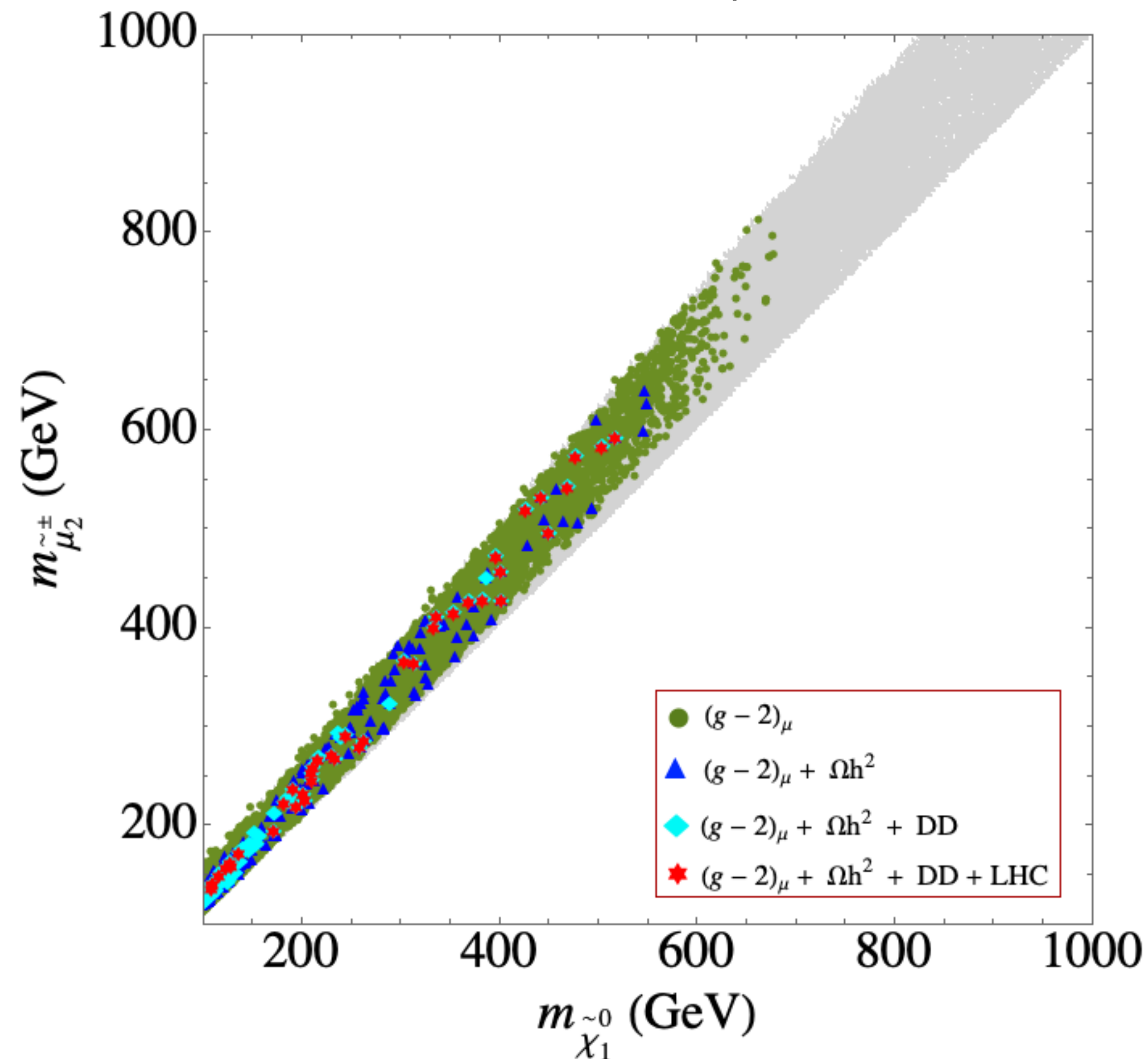


Reduced limit attributed by significant $\text{BR}(\tilde{\chi}_1^\pm \rightarrow \tilde{\tau}_1 \nu_\tau)$ and $\text{BR}(\tilde{\chi}_2^0 \rightarrow \tilde{\tau}_1 \tau)$, $\text{BR}(\tilde{\chi}_2^0 \rightarrow \tilde{\nu} \nu)$

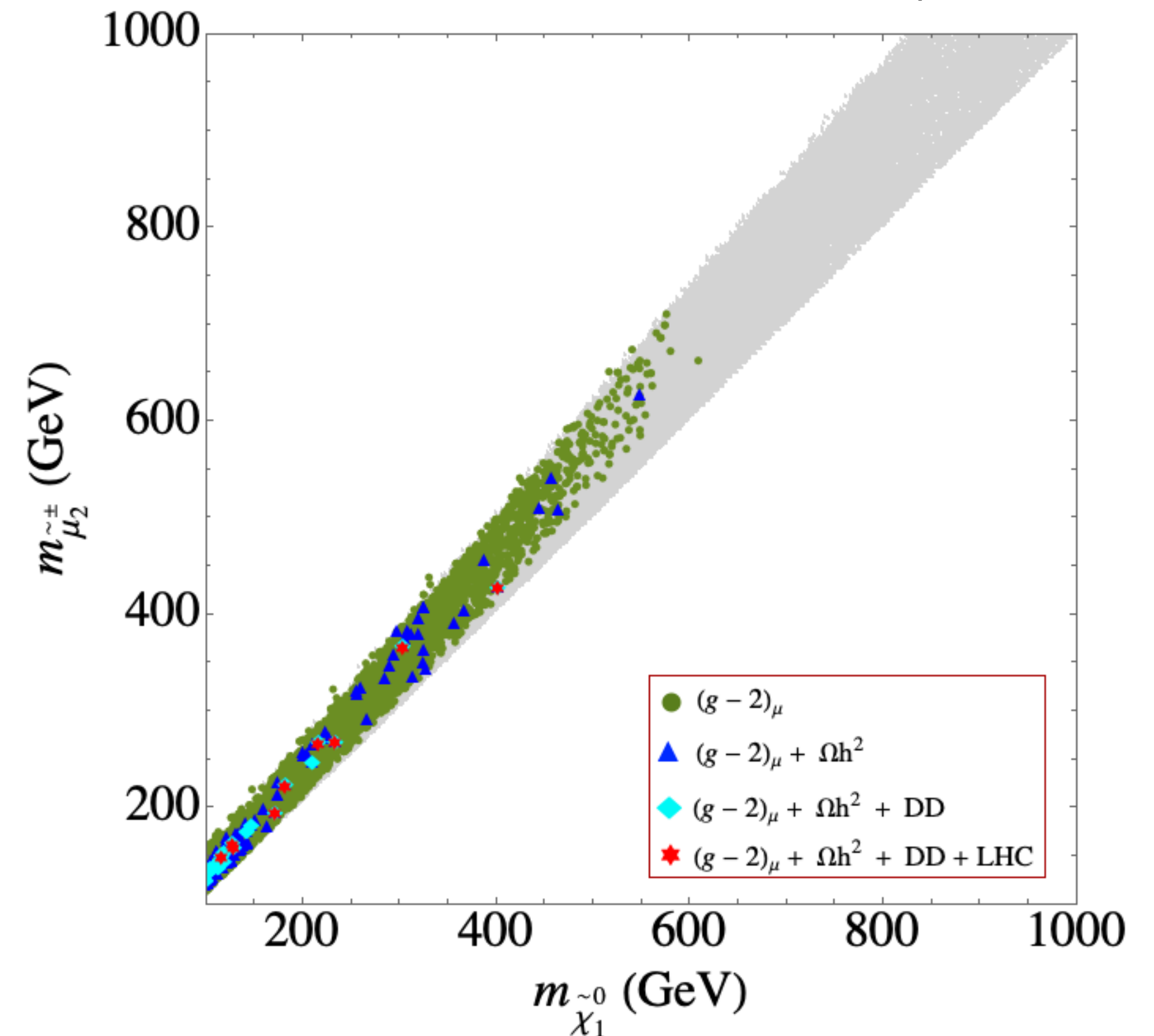
$(3l + \text{missing}ET)$ exclusion limit weakens

Slepton Co-annihilation: Case-R

Current $(g - 2)_\mu$ limit



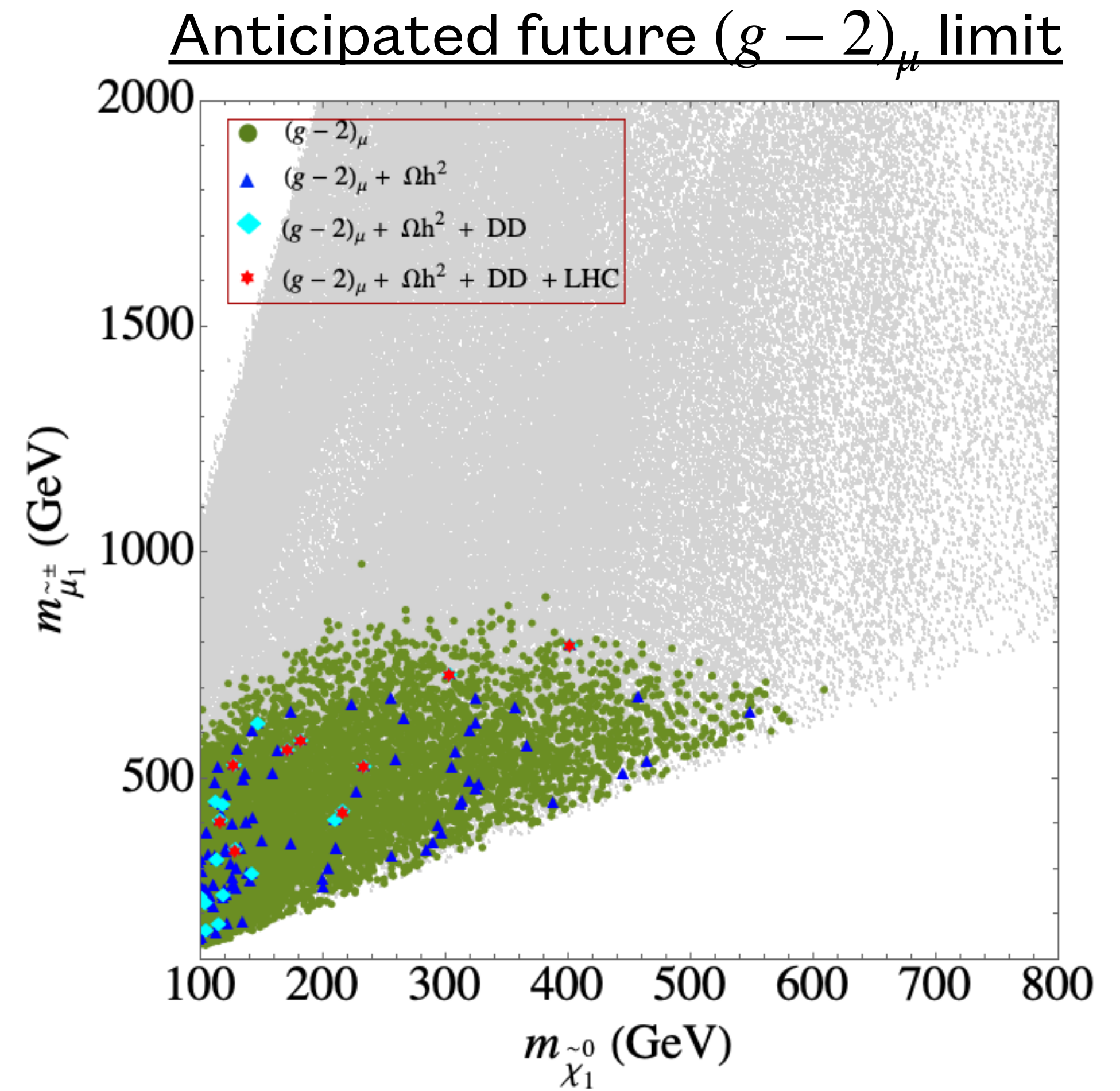
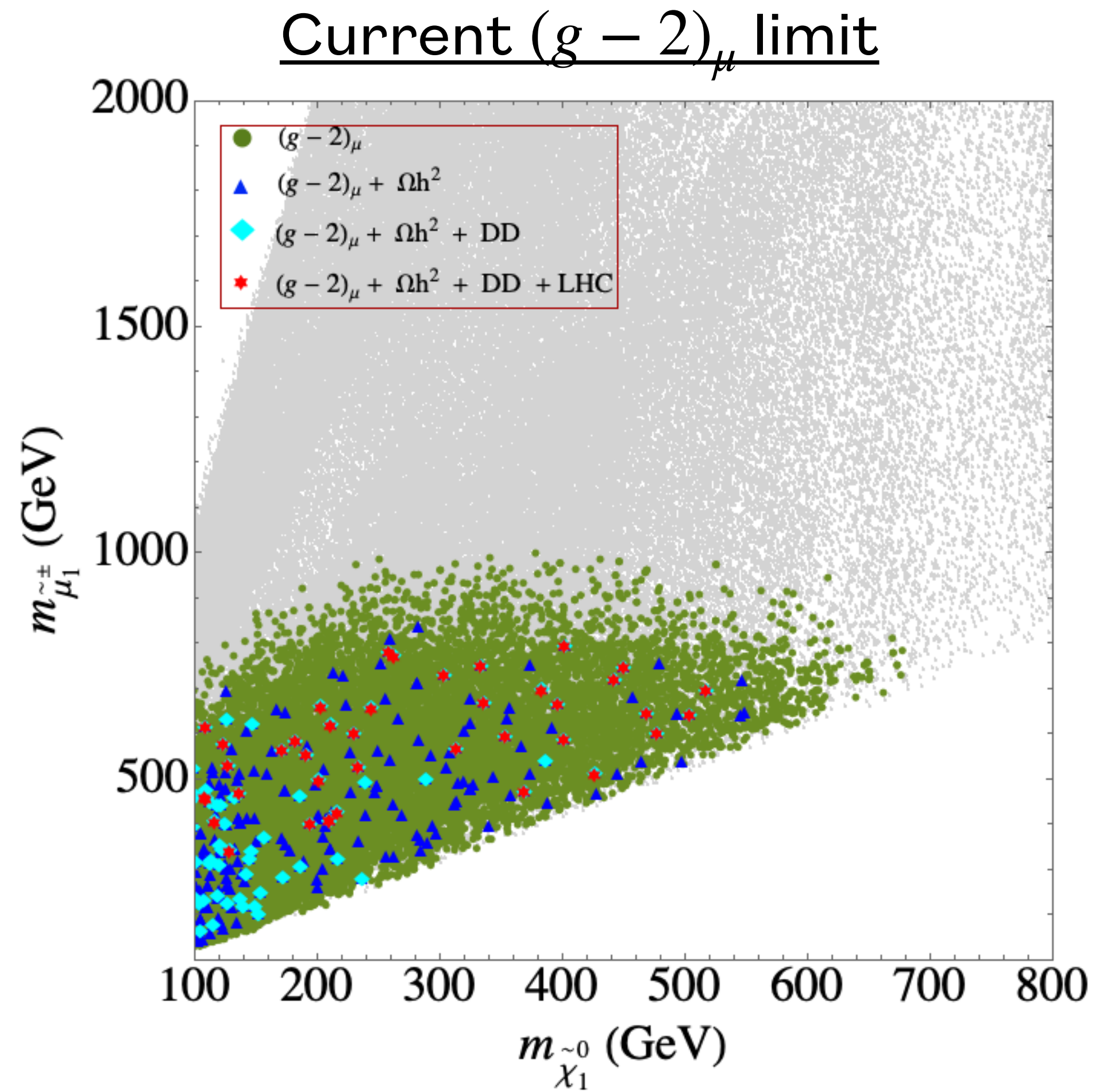
Anticipated future $(g - 2)_\mu$ limit



Right-sleptons are close in mass to LSP.

Small μ is favored, tension between DD and $(g - 2)_\mu$.

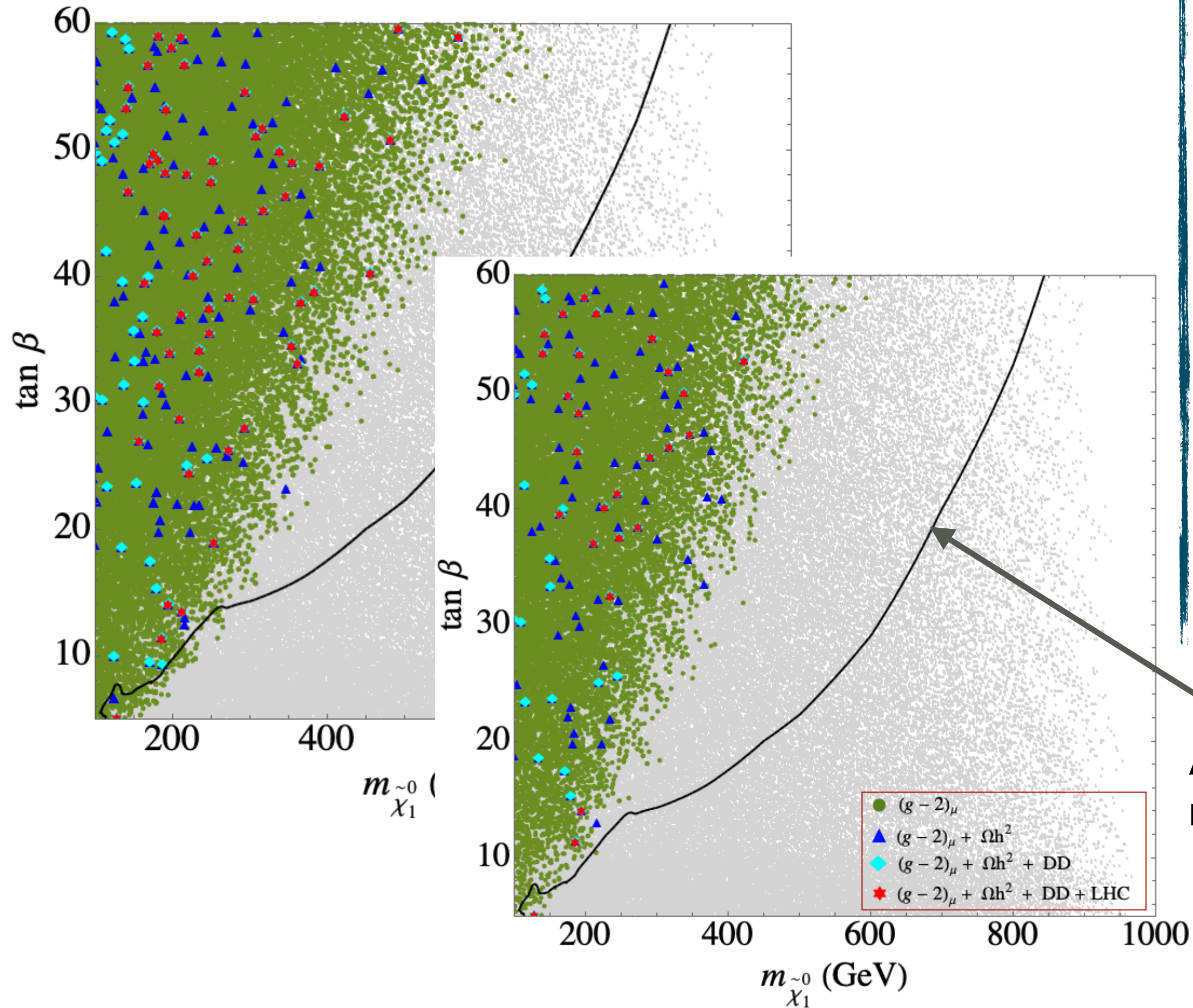
Slepton Co-annihilation: Case-R



Left-sleptons can not be too heavy to have relevant contribution to $(g - 2)_\mu$.

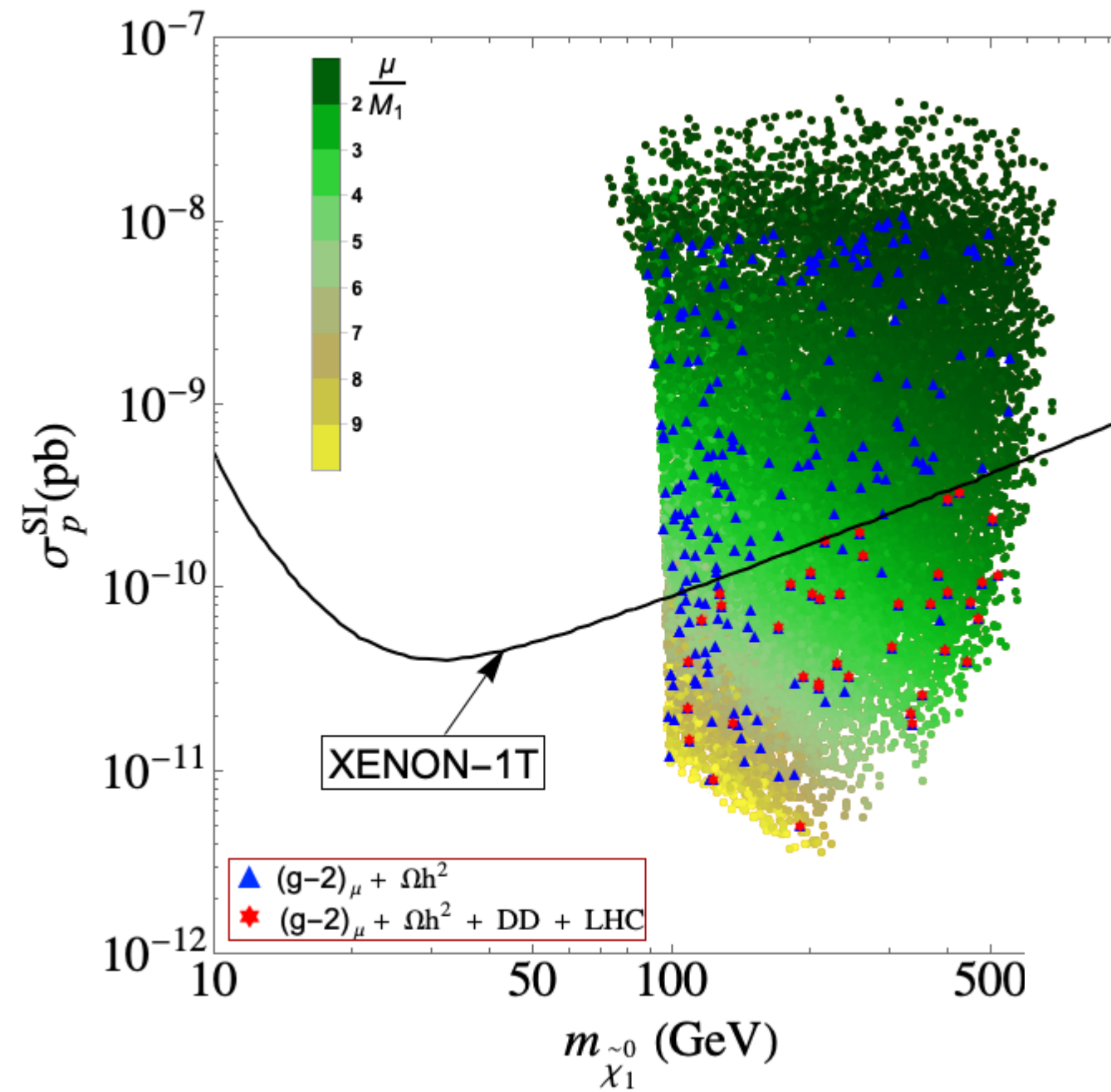
Get stringent constraint from LHC.

Further Comments

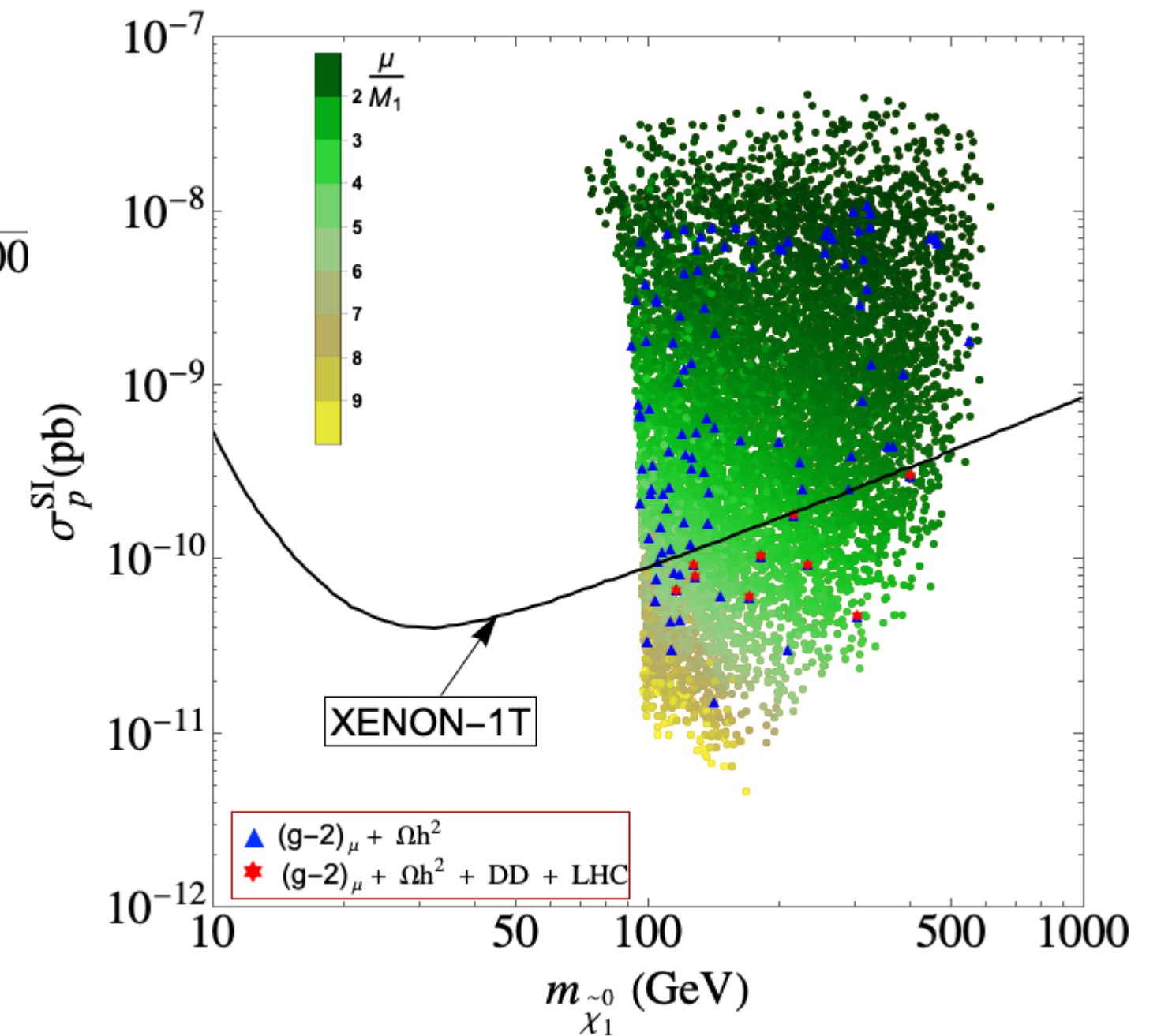


Chargino co-annihilation

A-pole annihilation
restricted



Slepton Co-annihilation
Case-R



Larger $\tan \beta$ can easily satisfy $(g-2)_\mu$.

Tension between DD and $(g-2)_\mu$.

Lowest and highest LSP : chargino coannihilation

Lowest LSP in Current (g-2)

Highest LSP in Current (g-2)

Highest LSP in Future (g-2)

Sample points	C1	C2	C3	Sample points	C1	C2	C3
M_1	133	579	430	BR($\tilde{\chi}_2^0 \rightarrow \tilde{\tau}_1 \tau$)	100	100	100
M_2	144	583	444				
μ	1329	1081	1024				
$\tan \beta$	5.1	59	52.7	BR($\tilde{\chi}_1^\pm \rightarrow \tilde{\tau}_1 \nu_\tau$)	100	100	100
$m_{\tilde{L}} = m_{\tilde{R}}$	170	678	540				
$m_{\tilde{\chi}_1^0}$	129	570	423				
$m_{\tilde{\chi}_2^0}$	150	605	464				
$m_{\tilde{\chi}_3^0}$	1338	1087	1032	BR($\tilde{e}_1 \rightarrow \tilde{\chi}_1^0 e$ $\rightarrow \tilde{\chi}_2^0 e$ $\rightarrow \tilde{\chi}_1^\pm \nu_e$)	20	14	16.4
$m_{\tilde{\chi}_4^0} \sim m_{\tilde{\chi}_2^\pm}$	1341	1093	1036		28	30	28.9
$m_{\tilde{\chi}_1^\pm}$	150	605	464		52	55	54.6
$m_{\tilde{e}_1, \tilde{\mu}_1}$	176	680	542				
$m_{\tilde{e}_2, \tilde{\mu}_2}$	176	680	541				
$m_{\tilde{\tau}_1}$	140	582	437				
$m_{\tilde{\tau}_2}$	205	765	629	BR($\tilde{e}_2 \rightarrow \tilde{\chi}_1^0 e$ $\rightarrow \tilde{\chi}_2^0 e$)	99.9	99.7	99.9
$m_{\tilde{\nu}}$	159	675	536		0.1	0.3	0.1
$\Omega_{\tilde{\chi}} h^2$	0.118	0.121	0.118				
$a_\mu^{\text{SUSY}} \times 10^{10}$	21.1	15.6	20.14				
$\sigma_p^{\text{SI}} \times 10^{10}$	0.39	2.3	1.12				

Searches for τ rich final states will be beneficial for further study.

BR in %

Lowest and highest LSP : Slepton coannihilation

Sample points	L1	L2	L3	Sample points	L1	L2	L3
M_1	131	541	508	BR($\tilde{\chi}_2^0 \rightarrow \tilde{l}_1 l$ $\rightarrow \tilde{\tau}_1 \tau$ $\rightarrow \tilde{\nu} \nu$ $\rightarrow \tilde{\chi}_1^0 h$ $\rightarrow \tilde{\chi}_1^0 Z$)	32	32.4	28
M_2	838	793	515		17	18.4	17.4
μ	720	1365	1012		34.5	49.2	54.6
$\tan \beta$	6.95	56.7	56		13	-	-
$m_{\tilde{l}_L}$	149	548	509		3.43	-	-
$m_{\tilde{l}_R}$	1172	1278	2349				
$m_{\tilde{\chi}_1^0}$	126	533	499				
$m_{\tilde{\chi}_2^0}$	706	816	535				
$m_{\tilde{\chi}_3^0}$	731	1369	1019				
$m_{\tilde{\chi}_4^0} \sim m_{\tilde{\chi}_2^\pm}$	889	1374	1025	BR($\tilde{\chi}_1^\pm \rightarrow \tilde{\nu}_l l$ $\rightarrow \tilde{\nu}_{\tau_1} \tau$ $\rightarrow \tilde{l}_1 \nu_l$ $\rightarrow \tilde{\tau}_1 \nu_\tau$ $\rightarrow W \tilde{\chi}_1^0$)	32	33.2	39.4
$m_{\tilde{\chi}_1^\pm}$	706	816	535		17	17	20.4
$m_{\tilde{e}_1, \tilde{\mu}_1}$	155	549	511		23.2	31.8	25.2
$m_{\tilde{e}_2, \tilde{\mu}_2}$	1173	1279	2349		11.7	17.7	15
$m_{\tilde{\tau}_1}$	155	534	509		16	-	-
$m_{\tilde{\tau}_2}$	1173	1286	2350				
$m_{\tilde{\nu}}$	135	544	505				
$\Omega_{\tilde{\chi}} h^2$	0.119	0.121	0.12	BR($\tilde{e}_1 \rightarrow \tilde{\chi}_1^0 e$)	100	100	100
$a_\mu^{\text{SUSY}} \times 10^{10}$	19.7	14.06	21.1	BR($\tilde{e}_2 \rightarrow \tilde{\chi}_1^0 e$)	100	100	99.2
$\sigma_p^{\text{SI}} \times 10^{10}$	0.8	0.46	2.13	$\rightarrow \tilde{\chi}_2^0 e$	-	-	0.5

Case-L

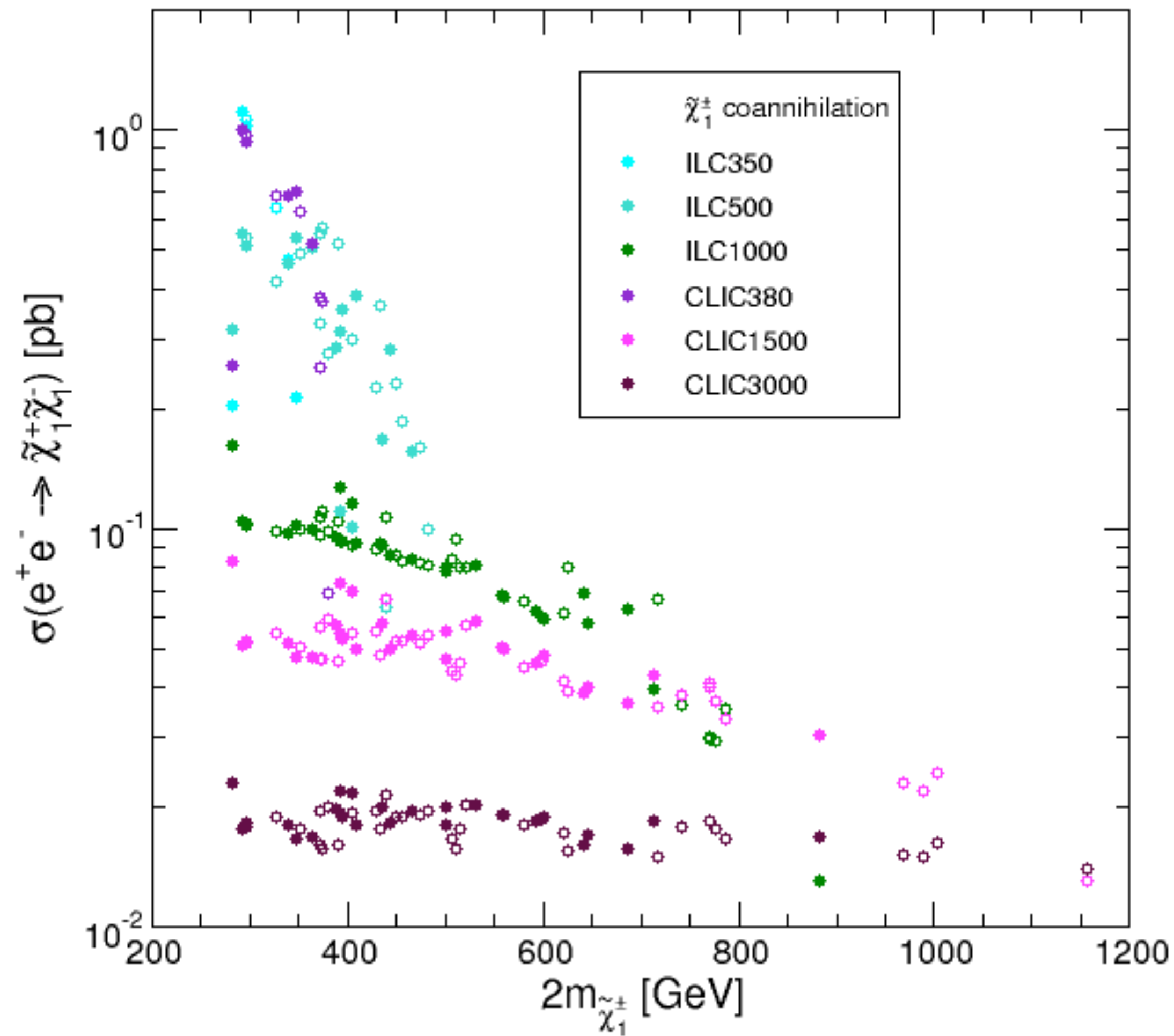
Sample points	R1	R2	R3	Sample points	R1	R2	R3
M_1	111	525	408	BR($\tilde{\chi}_2^0 \rightarrow \tilde{l}_2 l$ $\rightarrow \tilde{\tau}_2 \tau$ $\rightarrow \tilde{\chi}_1^0 h$ $\rightarrow \tilde{\chi}_1^0 Z$)	0.72	-	2.4
M_2	352	662	429		93.7	96.8	97.6
μ	812	1091	822		4.5	2.92	-
$\tan \beta$	20.5	58.5	59		0.99	-	-
$m_{\tilde{l}_L}$	458	695	794				
$m_{\tilde{l}_R}$	128	591	425				
$m_{\tilde{\chi}_1^0}$	109	518	402				
$m_{\tilde{\chi}_2^0}$	367	685	448	BR($\tilde{\chi}_1^\pm \rightarrow \tilde{l}_1 \nu_l$ $\rightarrow \tilde{\tau}_2 \nu_\tau$ $\rightarrow W \tilde{\chi}_1^0$)	-	-	-
$m_{\tilde{\chi}_3^0}$	823	1098	830		94.3	97	100
$m_{\tilde{\chi}_4^0} \sim m_{\tilde{\chi}_2^\pm}$	828	1105	838		5.7	2.8	-
$m_{\tilde{\chi}_1^\pm}$	367	685	448				
$m_{\tilde{e}_1, \tilde{\mu}_1}$	460	696	795				
$m_{\tilde{e}_2, \tilde{\mu}_2}$	136	592	428	BR($\tilde{e}_1 \rightarrow \tilde{\chi}_1^0 e$ $\rightarrow \tilde{\chi}_2^0 e$ $\rightarrow \tilde{\chi}_1^\pm \nu_e$)	42	95	9.2
$m_{\tilde{\tau}_2}$	119	526	406		19.6	1.7	32
$m_{\tilde{\tau}_1}$	464	747	807		38.3	3.2	58.7
$m_{\tilde{\nu}}$	453	692	792				
$\Omega_{\tilde{\chi}} h^2$	0.121	0.121	0.121				
$a_\mu^{\text{SUSY}} \times 10^{10}$	17.5	14.8	17.8	BR($\tilde{e}_2 \rightarrow \tilde{\chi}_1^0 e$)	100	100	100
$\sigma_p^{\text{SI}} \times 10^{10}$	0.23	1.2	3.1				

Case-R

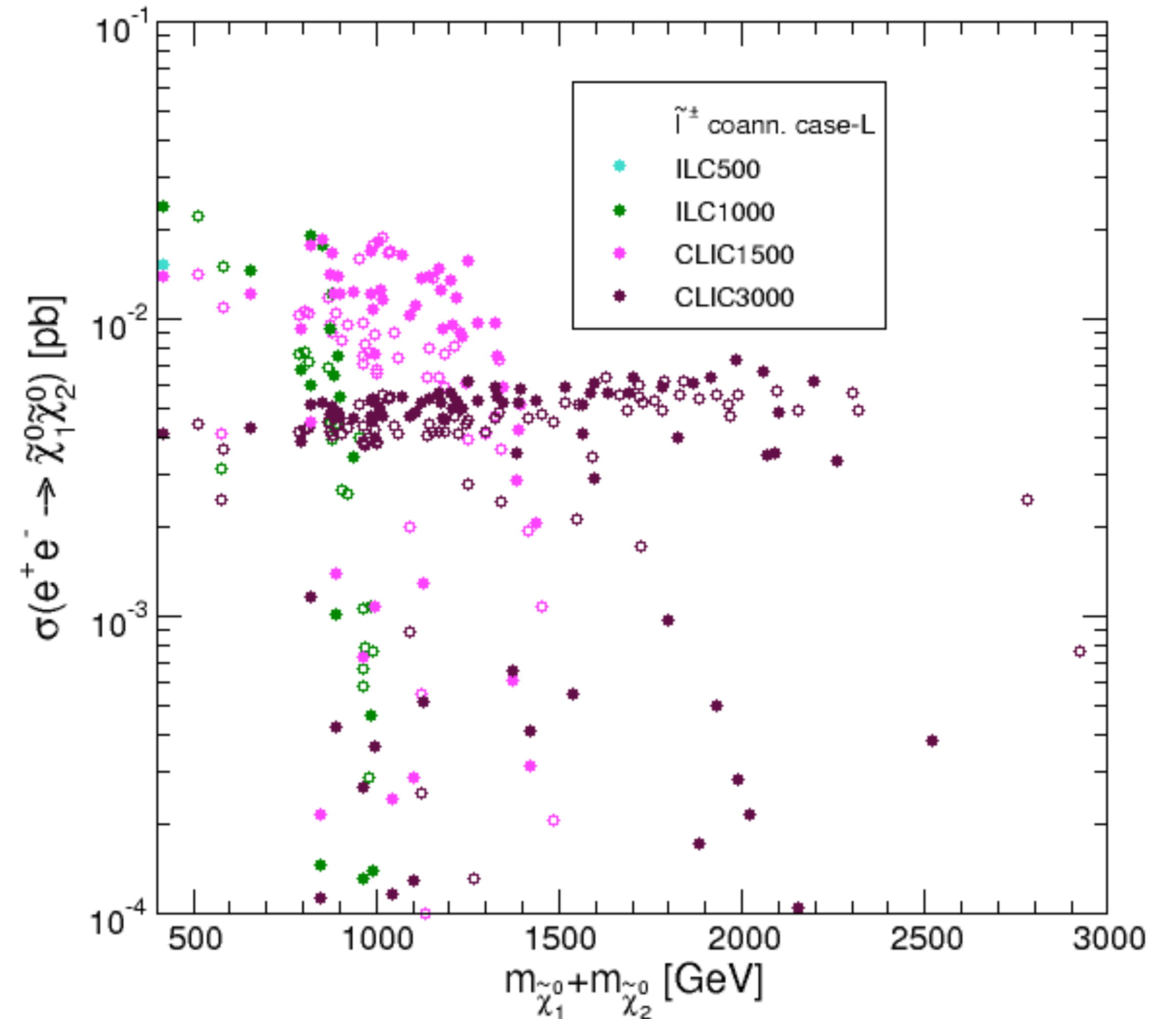
Lowest & Highest LSP mass in current (g-2) : L1, L2 (R1,R2)

Highest LSP mass in Future (g-2) : L3 (R3)

Target for future collider : ILC/CLIC

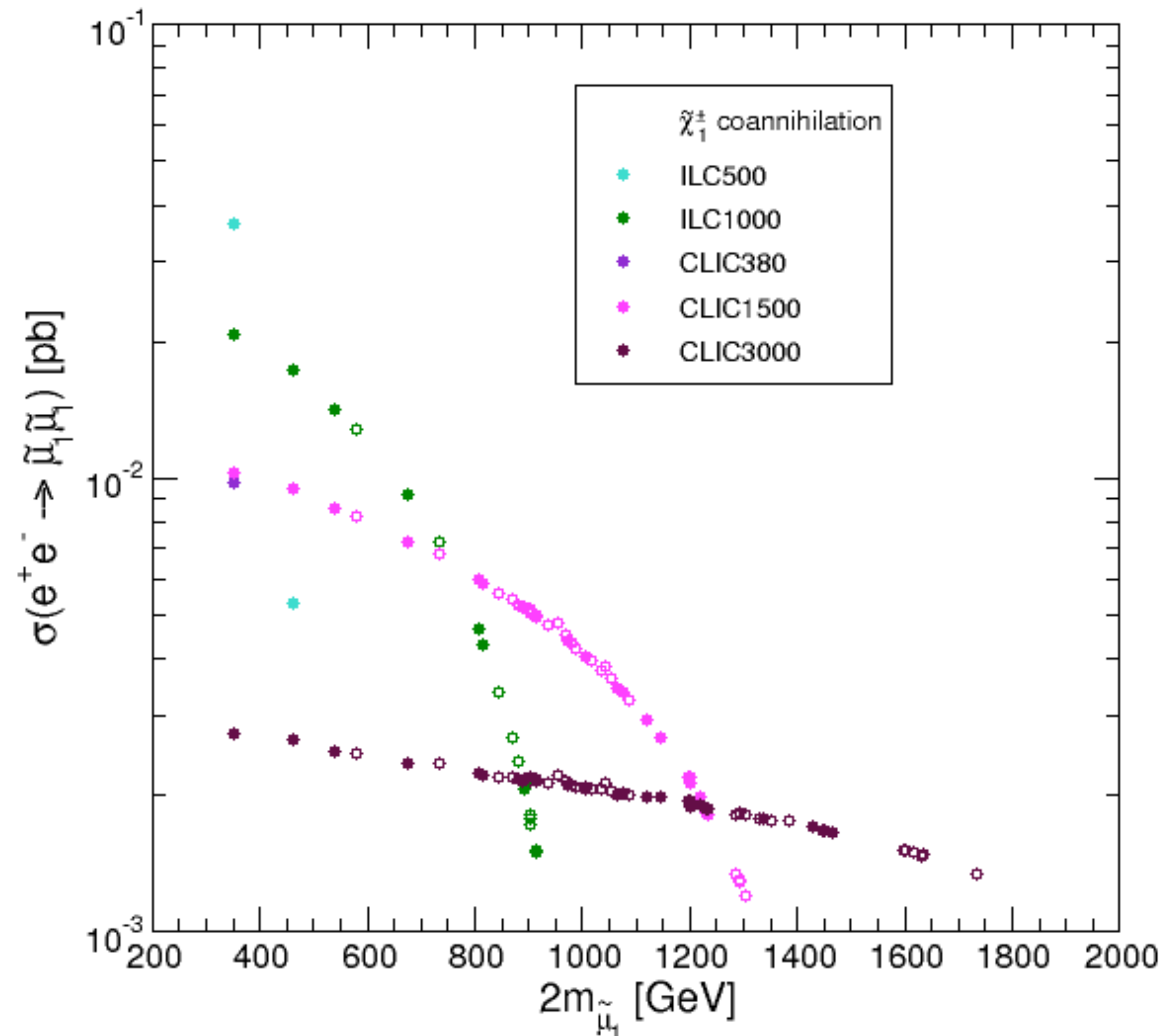


Chargino coannihilation

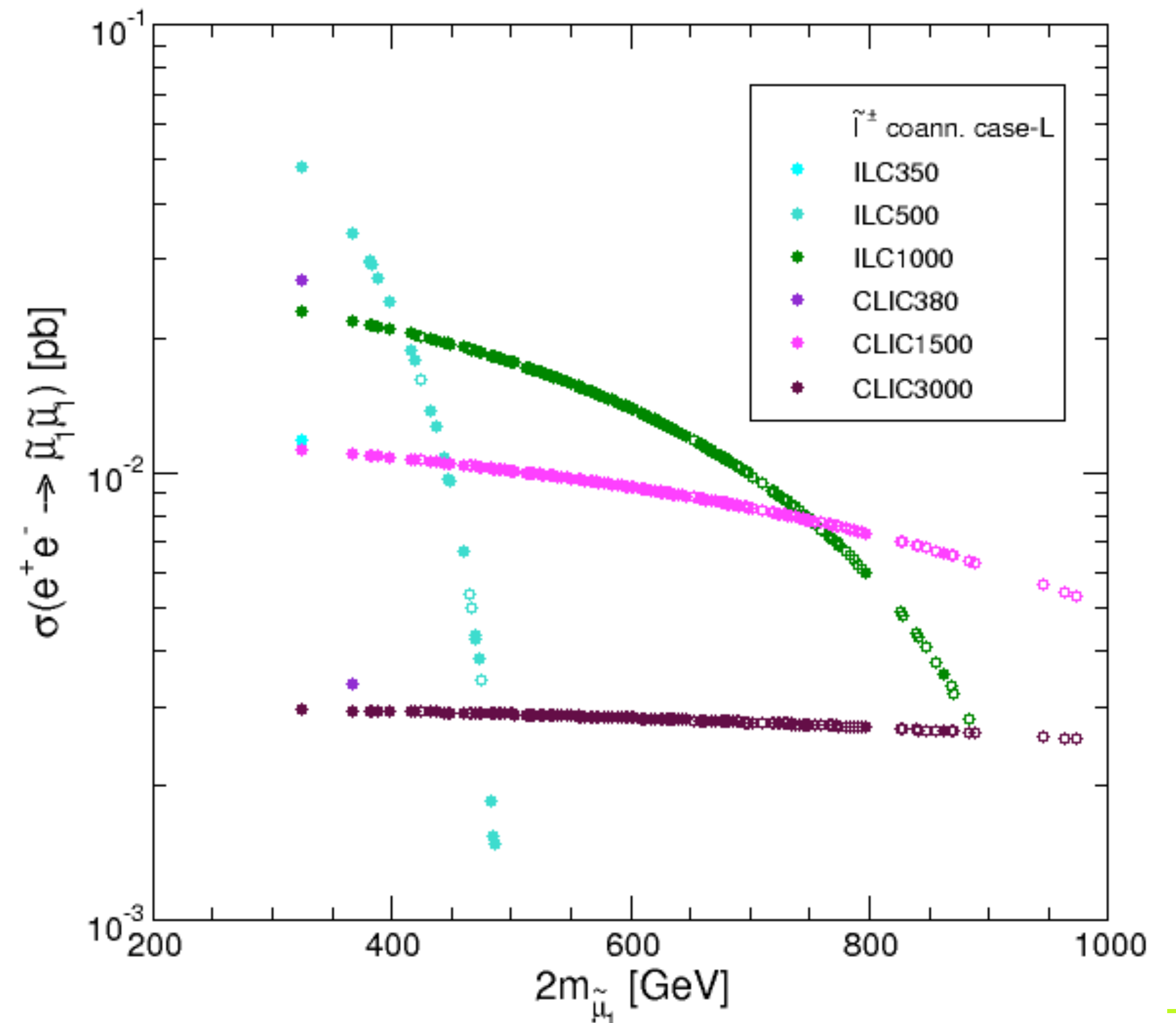


Slepton coannihilation : Case L

Target for future collider : ILC/CLIC



Chargino coannihilation



Slepton coannihilation : Case L

Conclusions

- Direct LHC bounds still have ample room for sub-TeV EW SUSY particles.
 - It is possible to constrain the EW MSSM with the help of indirect constraints along with the direct collider limits.
 - DM and muon ($g-2$) constraint put effective upper limit on EW SUSY masses while LHC limits restrict the mass ranges from below.
 - LHC exclusion bound strongly depends on EW gaugino composition. Proper recasting of ATLAS/CMS analysis relaxes the existing bound.
 - Searches for τ rich final states will be beneficial for further study.
 - Future colliders, HL-LHC, ILC/CLIC also have significant prospect for detection.
 - We await the new experiments results on muon ($g-2$) from Fermilab, J-PARC.
STAY TUNED!!!
-

THANK YOU!

BACKUP

$(g - 2)_\mu$

- Large discrepancy from the SM (more than 3σ):

$$a_\mu^{exp} - a_\mu^{SM} = (28.02 \pm 7.37) \times 10^{-10}.$$

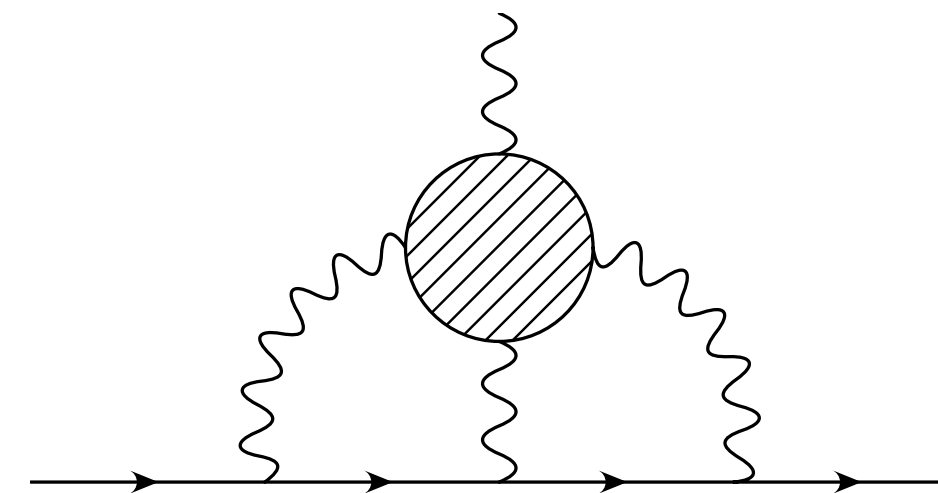
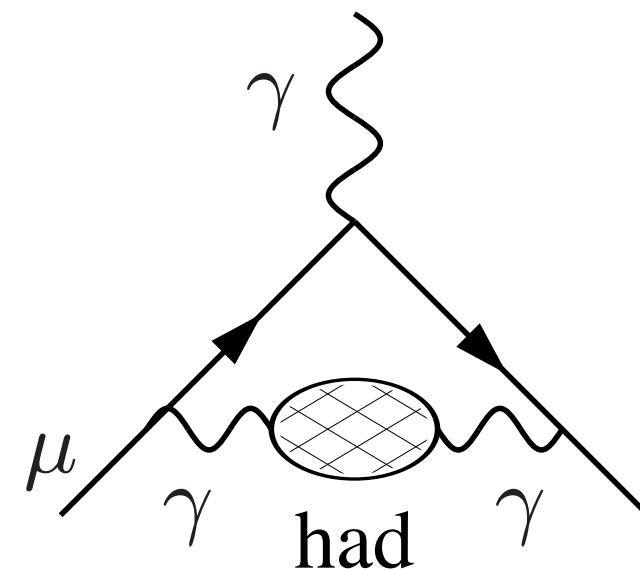
Keshavarzi, Nomura, Teubner '19

- Important probe for new physics. $\frac{\delta a_l}{a_l} \sim \frac{m_l^2}{\Lambda^2}$.
- SM contributions : QED, weak, hadronic vacuum polarization, hadronic light by light scattering.

- QED : complete calculation upto 5 loops. EW : two loops.

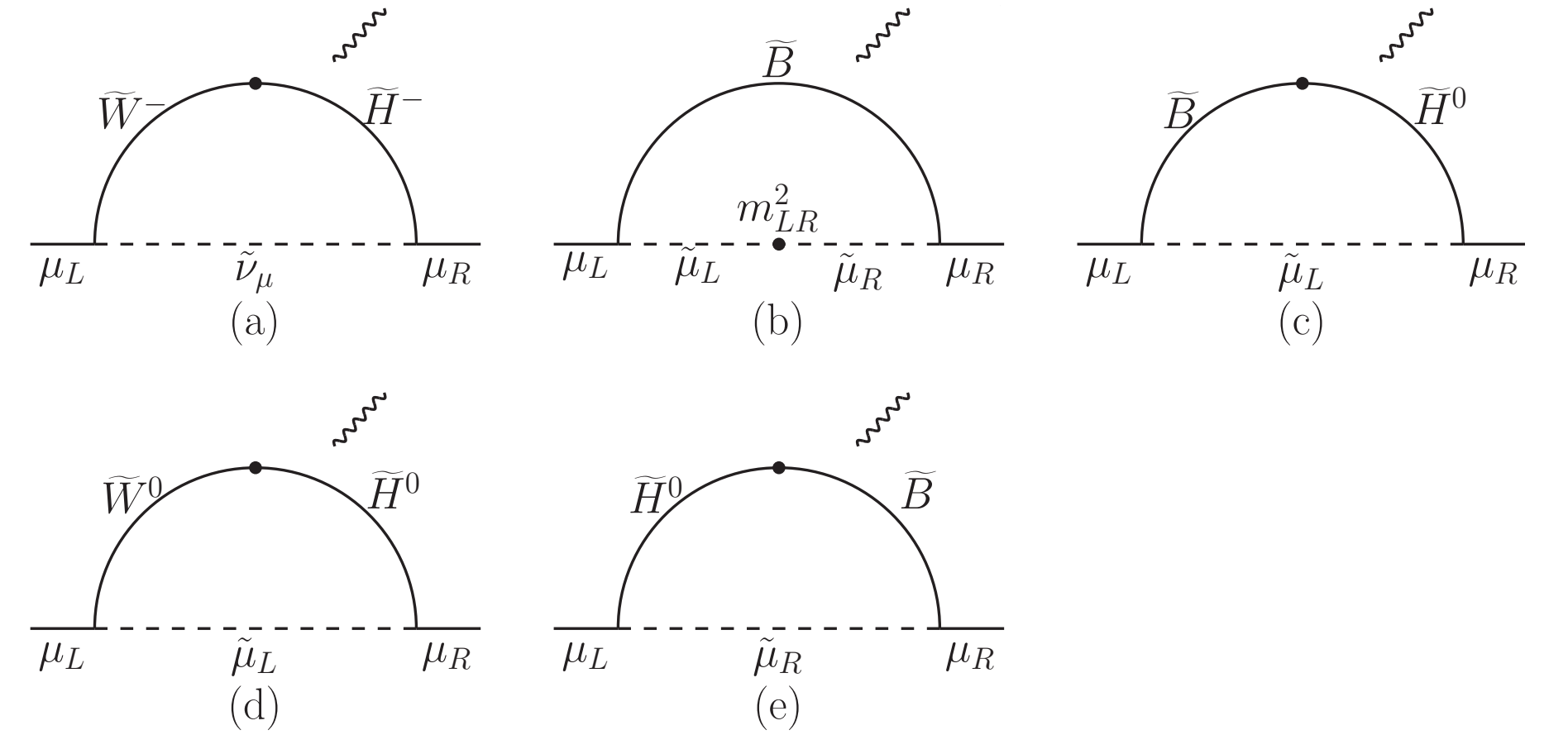
Aoyama, Hayakawa, Kinoshita, Nio '17, Ishikawa, Nakazawa, Yasu '18,
Heinemeyer, Stöckinger, Weiglein '04

- Uncertainty dominated by non-perturbative, hadronic sector.



SUSY contributions to $(g - 2)_\mu$

$$\begin{aligned} \Delta a_\mu(\tilde{W}, \tilde{H}, \tilde{\nu}_\mu) &\simeq 15 \times 10^{-9} \left(\frac{\tan \beta}{10} \right) \left(\frac{(100 \text{ GeV})^2}{M_{2\mu}} \right) \left(\frac{f_C}{1/2} \right), \\ \Delta a_\mu(\tilde{W}, \tilde{H}, \tilde{\mu}_L) &\simeq -2.5 \times 10^{-9} \left(\frac{\tan \beta}{10} \right) \left(\frac{(100 \text{ GeV})^2}{M_{2\mu}} \right) \left(\frac{f_N}{1/6} \right), \\ \Delta a_\mu(\tilde{B}, \tilde{H}, \tilde{\mu}_L) &\simeq 0.76 \times 10^{-9} \left(\frac{\tan \beta}{10} \right) \left(\frac{(100 \text{ GeV})^2}{M_{1\mu}} \right) \left(\frac{f_N}{1/6} \right), \\ \Delta a_\mu(\tilde{B}, \tilde{H}, \tilde{\mu}_R) &\simeq -1.5 \times 10^{-9} \left(\frac{\tan \beta}{10} \right) \left(\frac{(100 \text{ GeV})^2}{M_{1\mu}} \right) \left(\frac{f_N}{1/6} \right), \\ \Delta a_\mu(\tilde{\mu}_L, \tilde{\mu}_R, \tilde{B}) &\simeq 1.5 \times 10^{-9} \left(\frac{\tan \beta}{10} \right) \left(\frac{(100 \text{ GeV})^2}{m_{\tilde{\mu}_L}^2 m_{\tilde{\mu}_R}^2 / M_{1\mu}} \right) \left(\frac{f_N}{1/6} \right). \end{aligned}$$



Endo, Hamaguchi, Iwamoto, Yoshinaga'13